

Angular Momentum Presentation

Week 2

from Iskenderian and Salstein, *Monthly Weather Review* (1998)

2. Definitions and data sources

The global atmospheric angular momentum about the earth's axis can be expressed as

$$\text{AAM} = M_r + M_\Omega, \quad (1)$$

where

$$M_r = \frac{a^3}{g} \iiint u \cos^2 \phi \, d\phi \, d\lambda \, dp \quad (2)$$

is the entire atmosphere's angular momentum associated with its motion relative to the rotating solid earth, a is the earth's radius, g is acceleration due to gravity, u is zonal wind, and the integral is performed over all latitudes ϕ , longitudes λ , and pressures p . The angular mo-

mentum associated with the rotation of the atmosphere's mass is

$$M_{\Omega} = \frac{a^4 \Omega}{g} \iint p_s \cos^3 \phi \, d\phi \, d\lambda, \quad (3)$$

where Ω is the mean rotation rate of the earth and p_s the surface pressure (Rosen 1993).

The conservation of angular momentum states that changes in AAM are related to the surface torque by the relationship

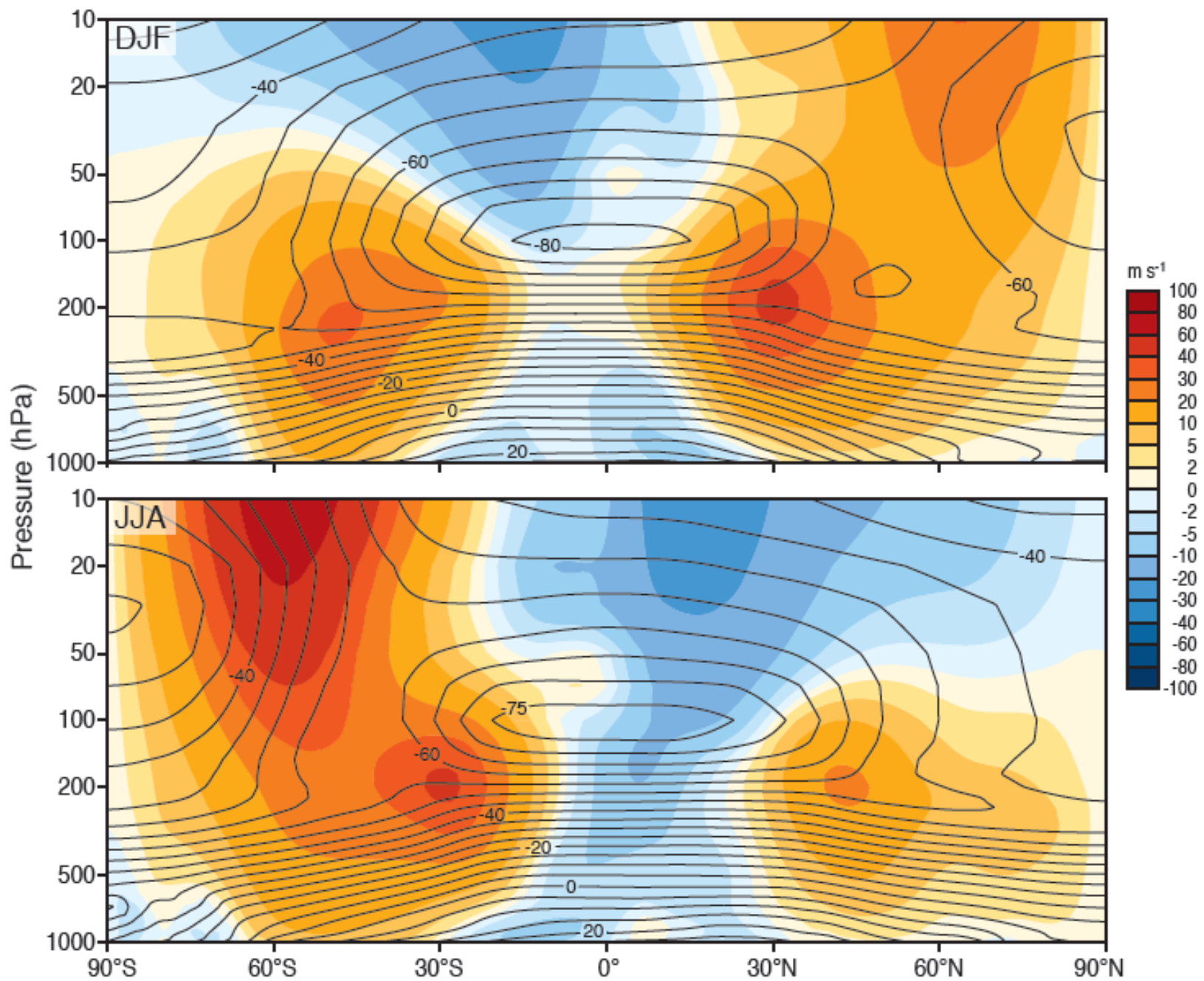
$$\frac{d}{dt} \text{AAM} = T_m + T_f, \quad (4)$$

where the mountain torque (T_m) and friction torque (T_f) are defined through the following relationships (White 1991):

$$T_m = -a^2 \iint p_s \frac{\partial H}{\partial \lambda} \cos \phi \, d\phi \, d\lambda, \quad (5)$$

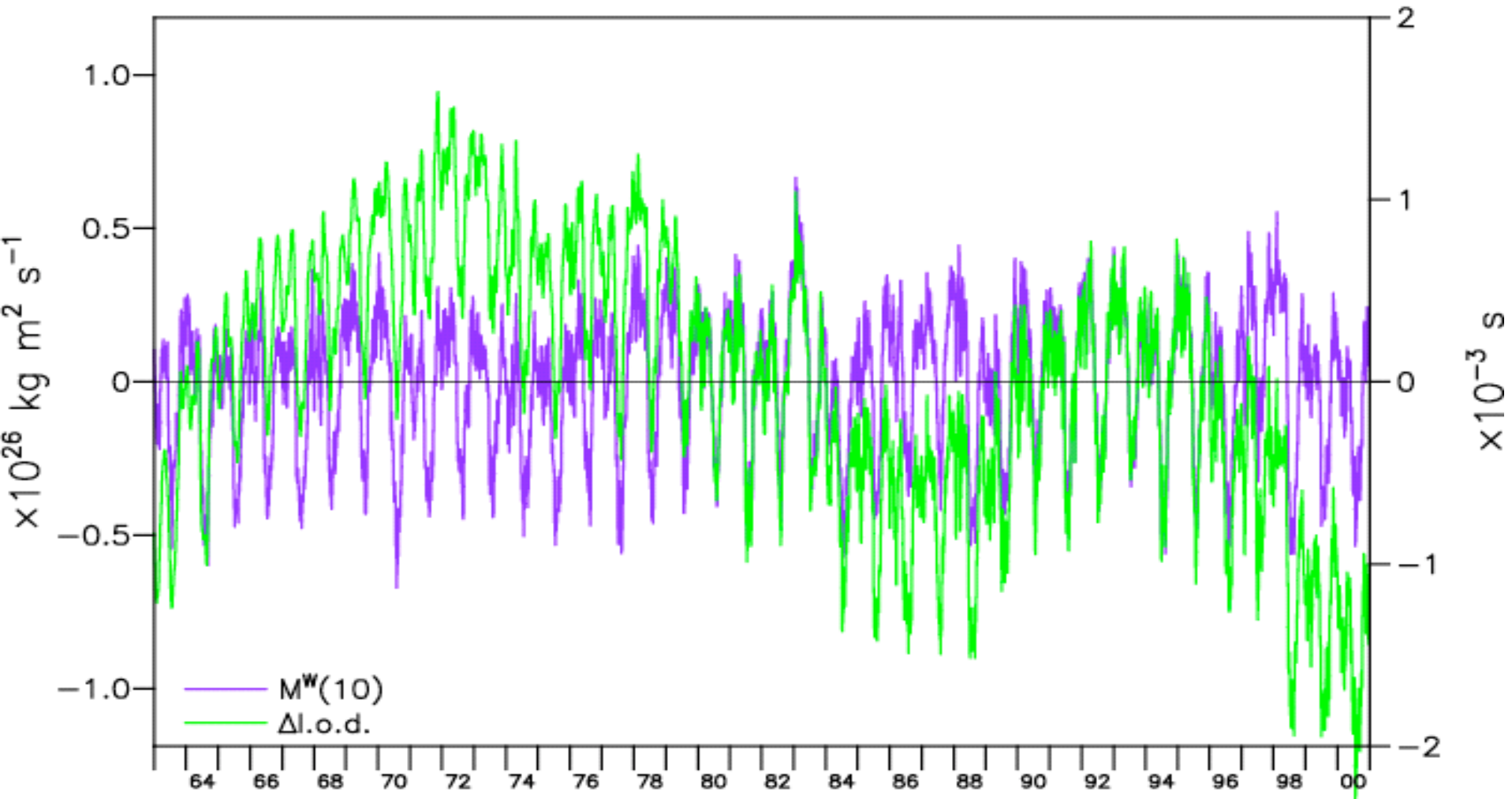
$$T_f = a^3 \iint \tau \cos^2 \phi \, d\phi \, d\lambda. \quad (6)$$

Here, τ is surface stress and H indicates the height of the sloping topography.



tein

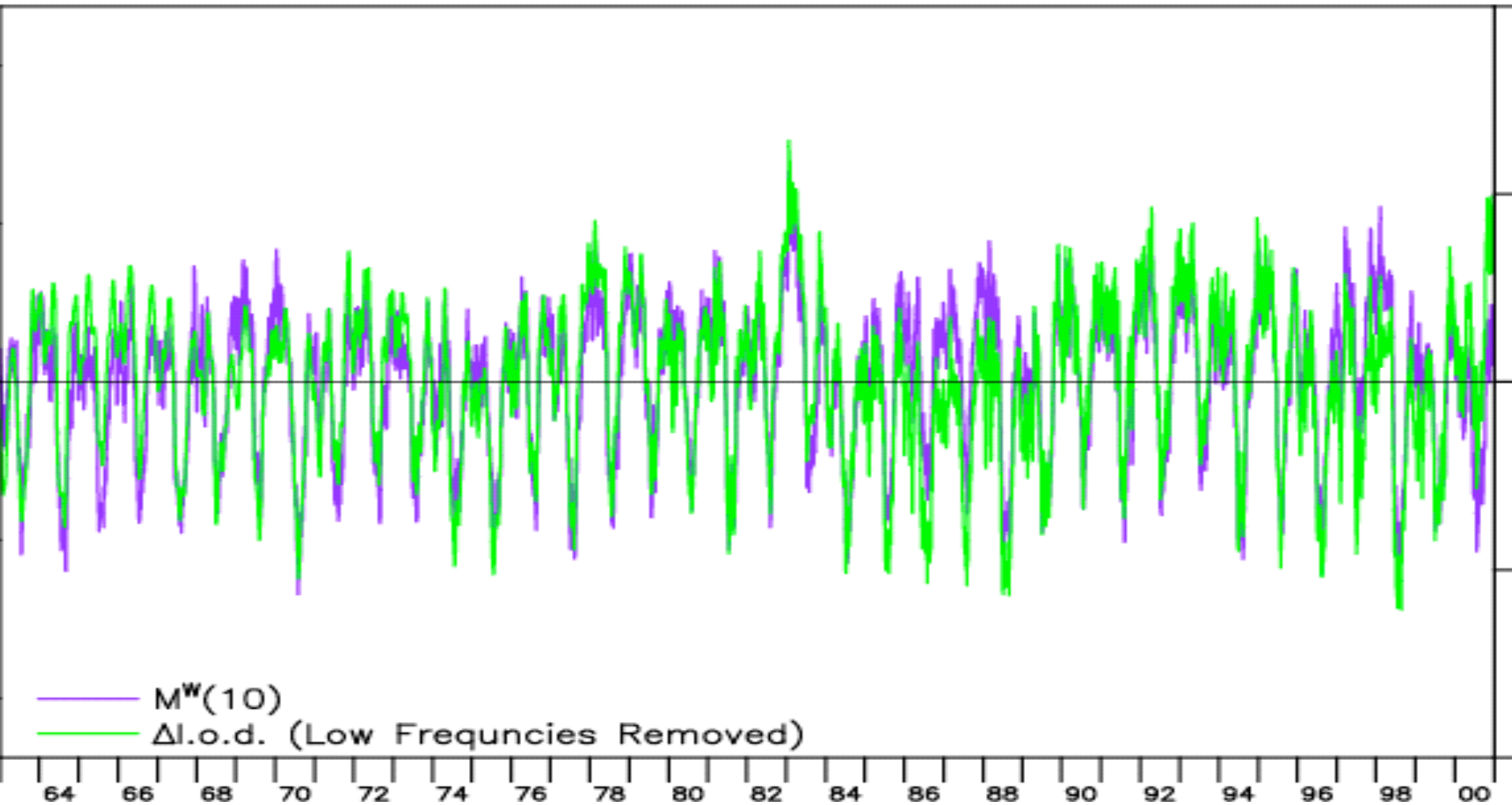
AAM and Length of Day



$$\Delta M = k \times \Delta(L.O.D.)$$

Courtesy of David Salstein

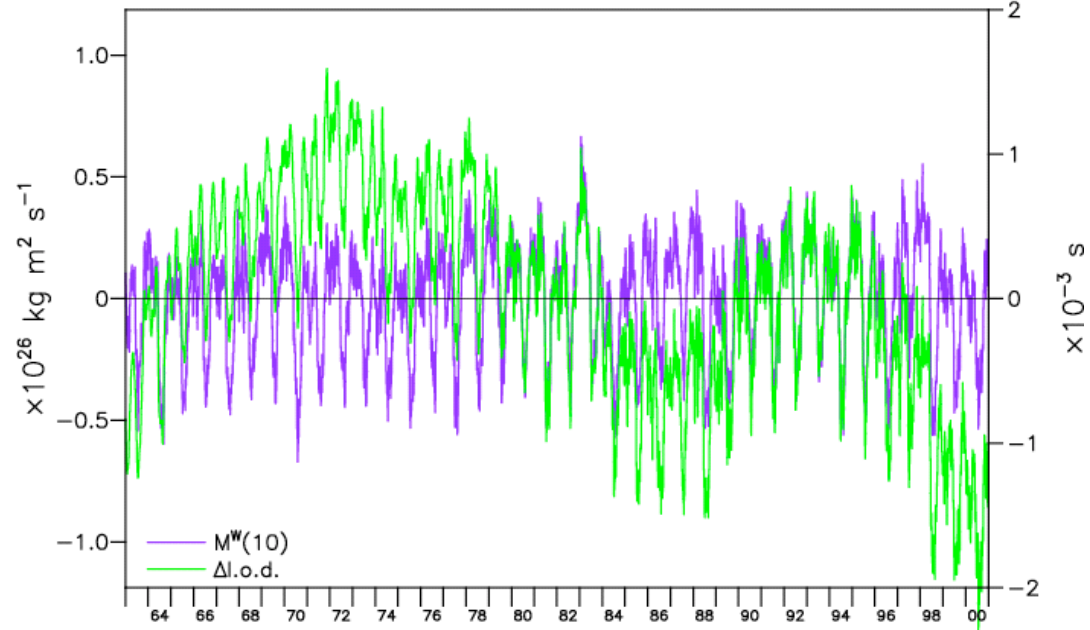
AAM and Length of Day



Filtered

Courtesy of David Salstein

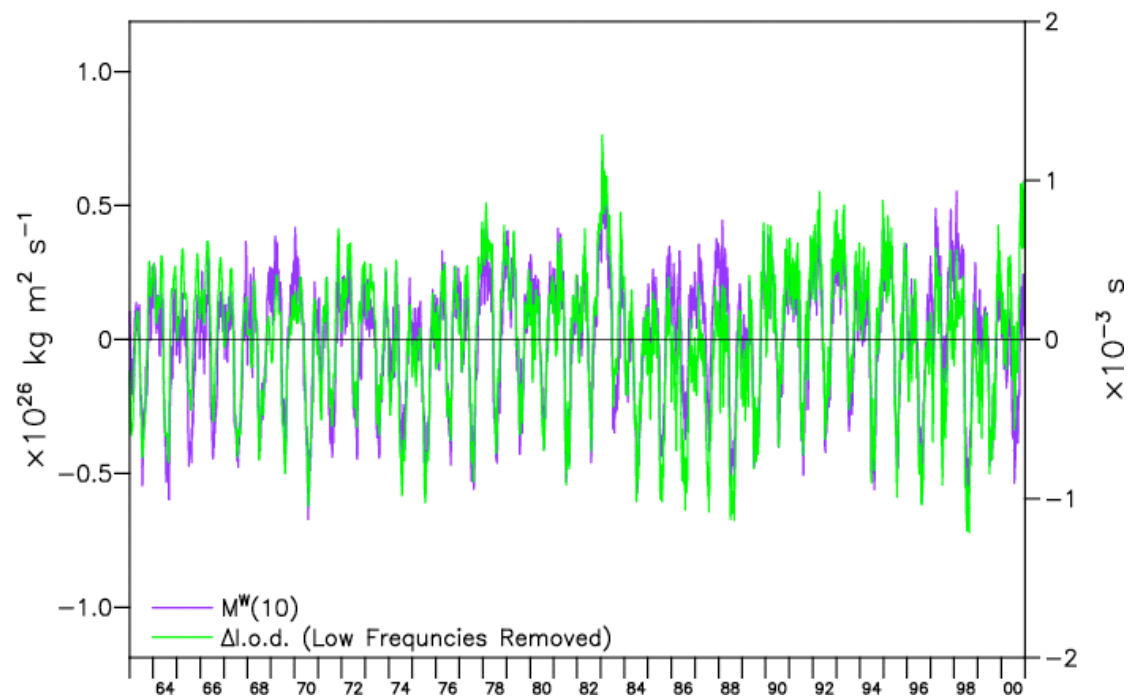
Raw



$$\Delta M = k \times \Delta(\text{L.O.D.})$$

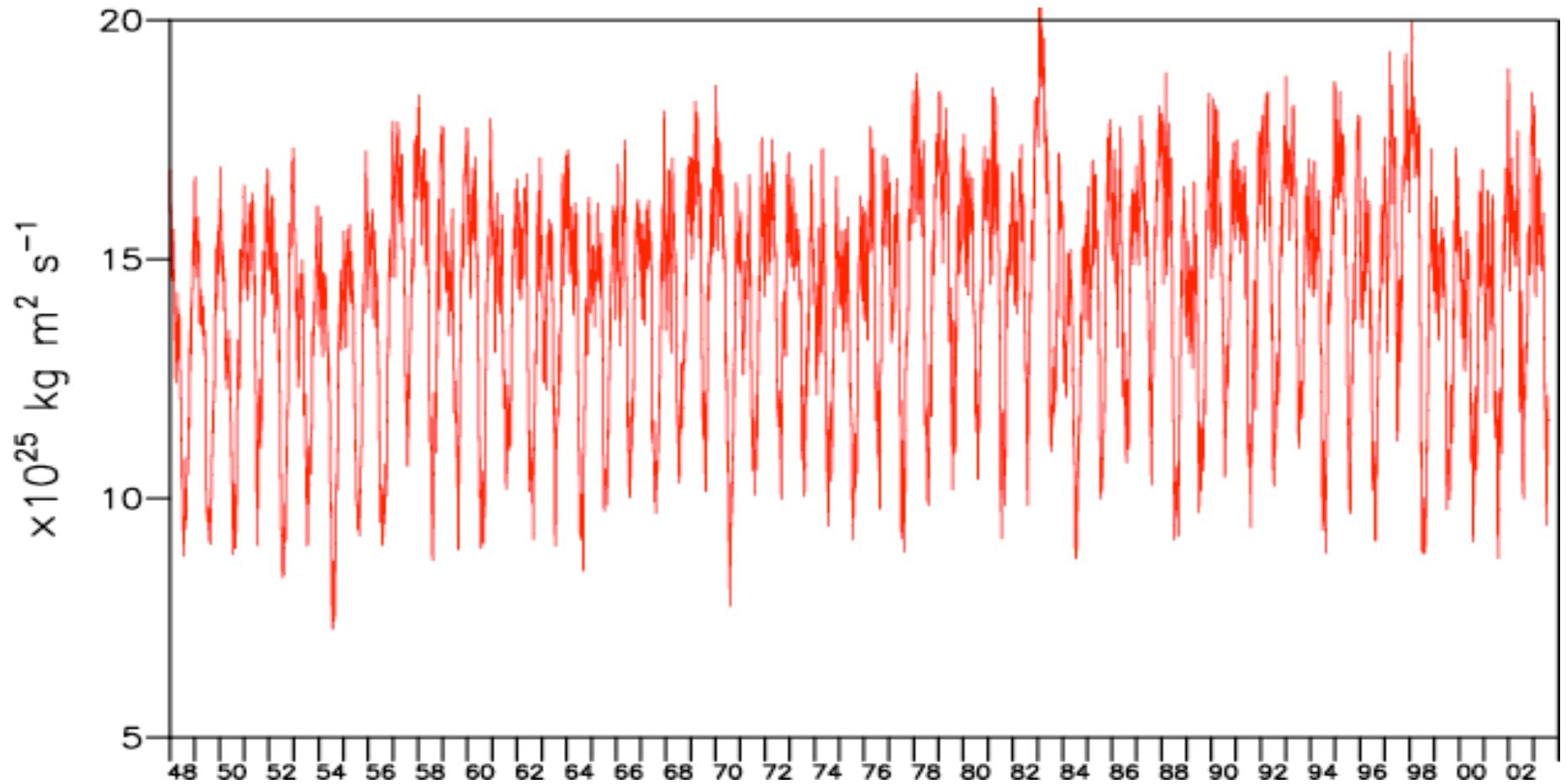
AAM and Length of Day

Filtered



Courtesy of David Salstein

NCEP/NCAR Reanalysis AAM



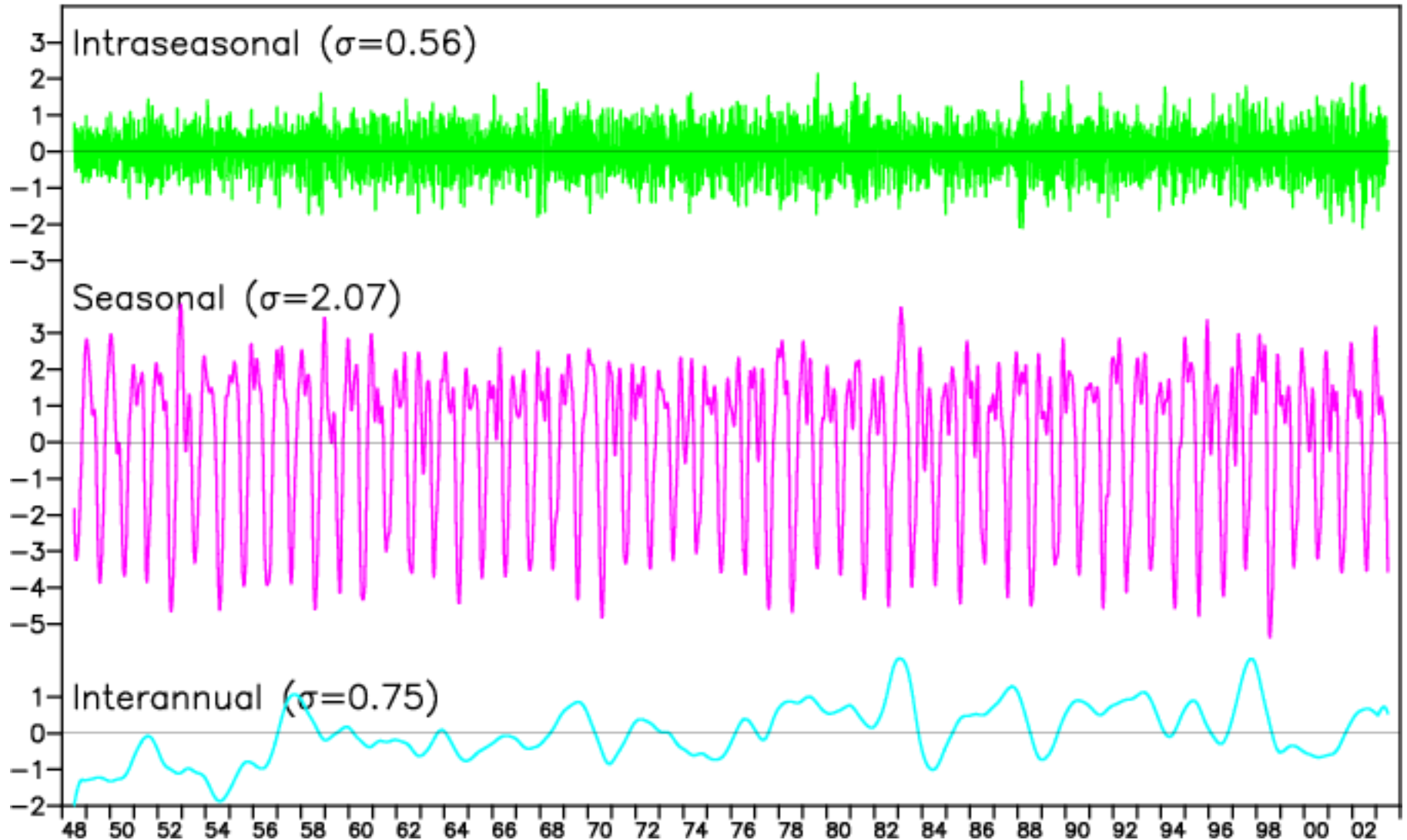
Note large seasonal cycle in angular momentum: factor of 2
between Jan/Feb and Jul/Aug

Courtesy of David Salstein

Breakdown into frequency bands

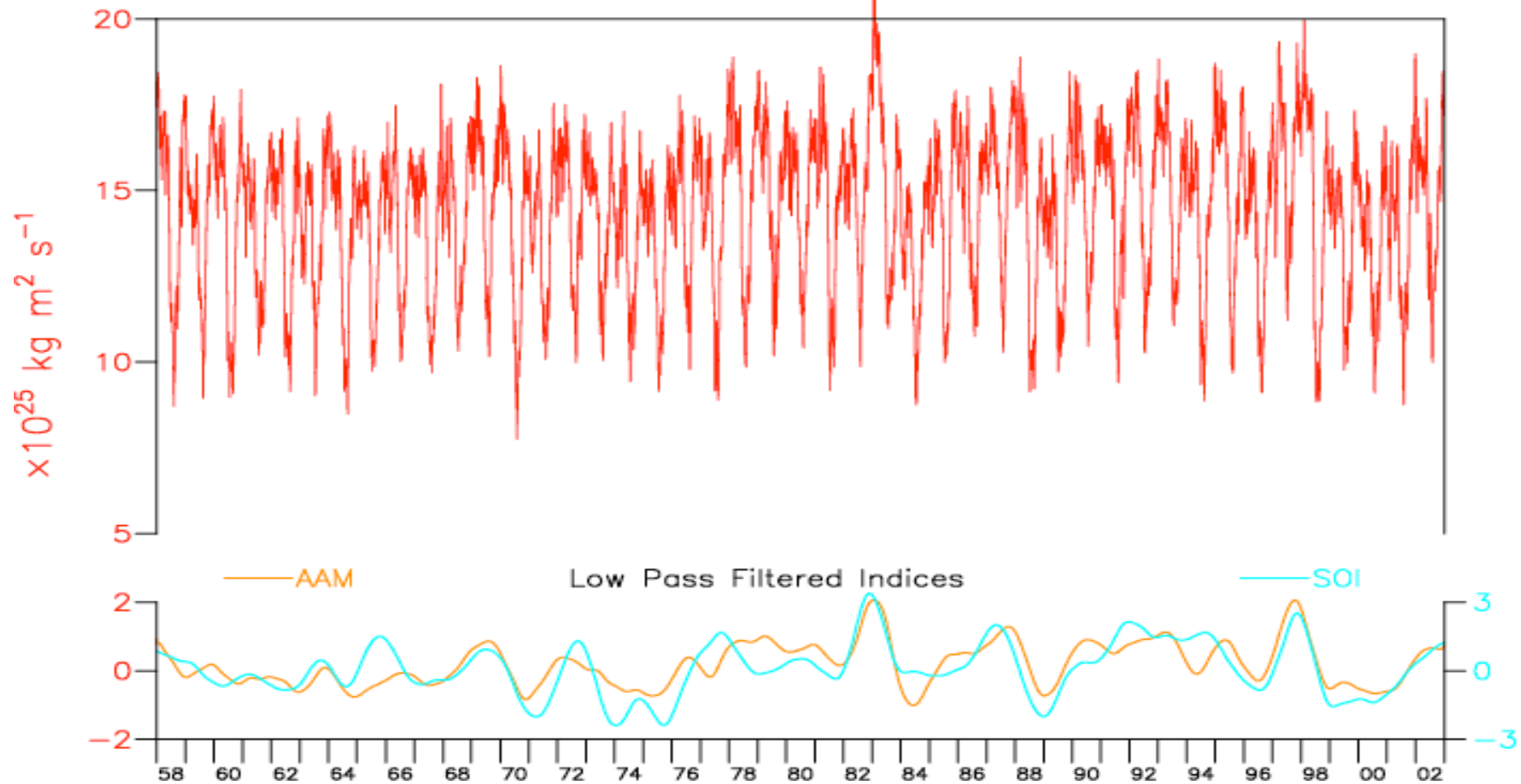
Jan. 1948–Dec. 2003

$\times 10^{25} \text{ kg m}^2 \text{ s}^{-1}$



Courtesy of David Salstein

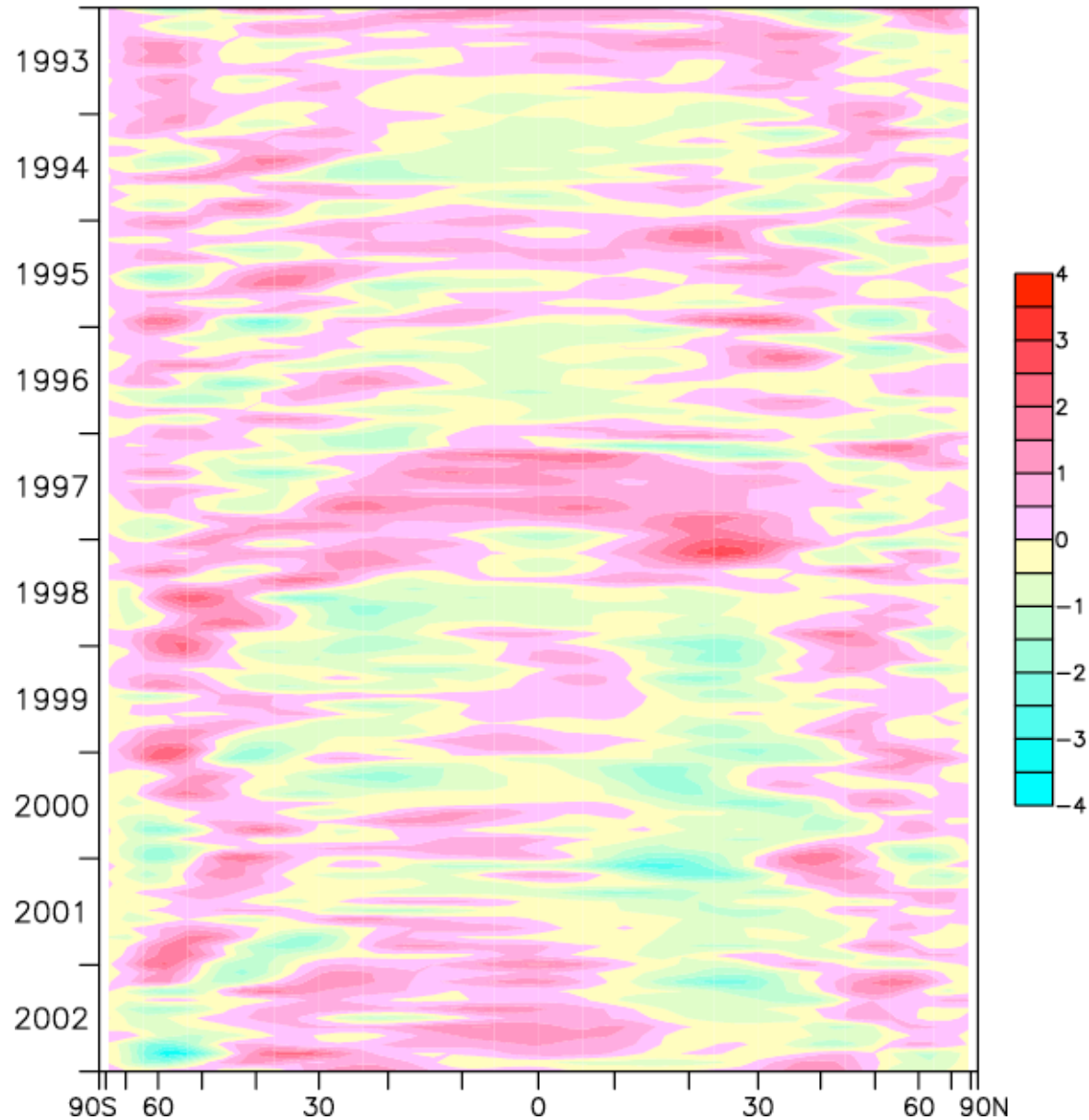
Atmospheric Angular Momentum (AAM), 1958–2002



SOI = -Southern Oscillation Index; peaks correspond to El Nino events

Courtesy of David Salstein

Reanalysis Belt $M^W(10)$ Anomaly
($\times 10^{24} \text{ kg m}^2 \text{ s}^{-1}$)



Positive AAM
anomalies
in tropics
coincide with
El Nino events
(e.g., 1997-98)

Courtesy of David Salstein

Mountain and friction torques

$$T_{mountain} = -R^2 \int \int p_s \frac{\partial H}{\partial \lambda} \cos \phi d\phi d\lambda$$

$$T_{friction} = R^3 \int \int \tau \cos^2 \phi d\phi d\lambda$$

$$T_{gravity-wave} = \begin{array}{l} \text{Frictional related to sub-grid scale} \\ \text{Action in the atmospheric model} \end{array}$$

R =Earth radius, $\mathbf{p}_s$ =surface pressure, $\mathbf{H}$ =topographic height

τ =frictional stress, related to winds and roughness (model)

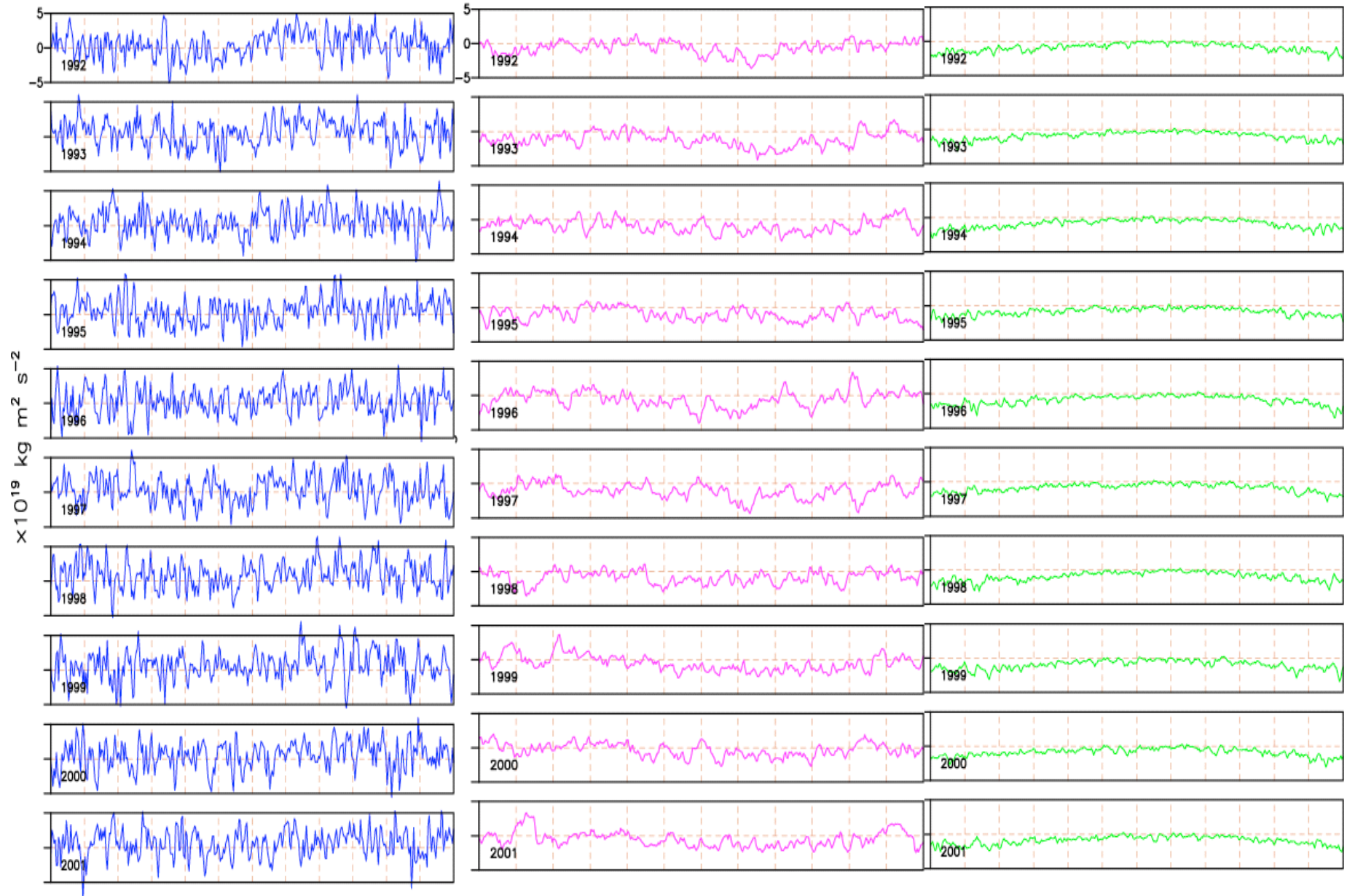
ϕ =longitude λ =latitude

Courtesy of David Salstein

Mountain Torque

Friction Torque

Gravity Wave Torque



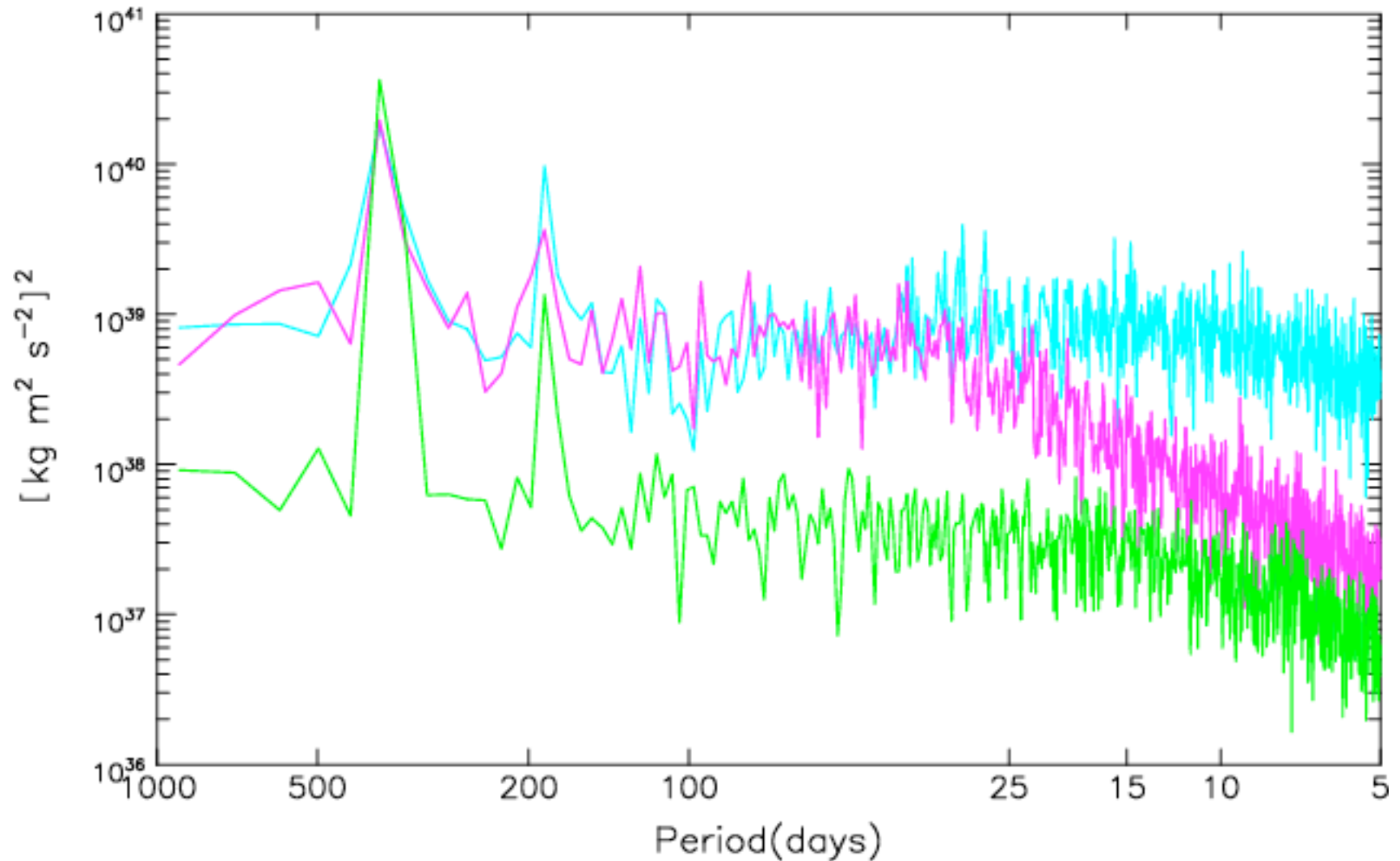
Courtesy of David Salstein

NCEP Reanalysis Torques Power Spectra (1958–2002)

Mountain

Friction

Gravity Wave



Courtesy of David Salstein

Regional Sources of Mountain Torque Variability and High-Frequency Fluctuations in Atmospheric Angular Momentum

HAIG ISKENDERIAN AND DAVID A. SALSTEIN

MONTHLY WEATHER REVIEW

VOLUME 126

1682

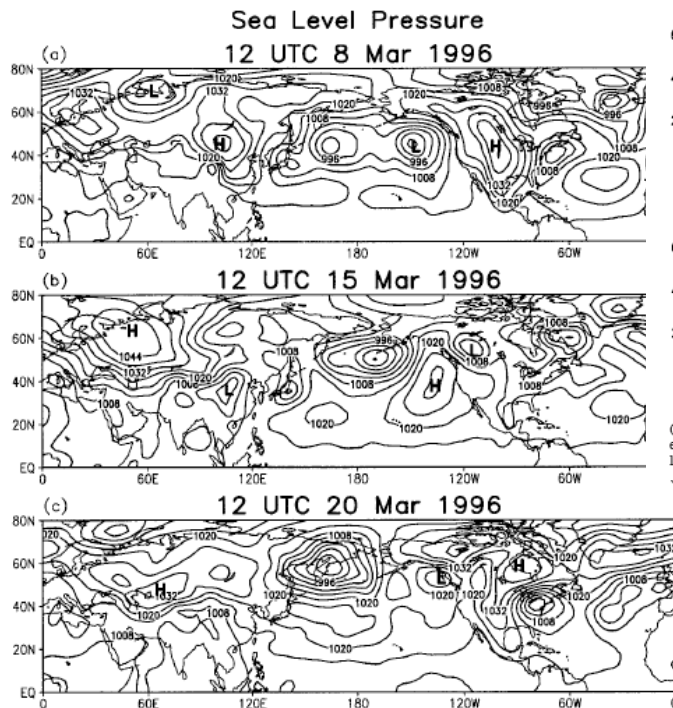
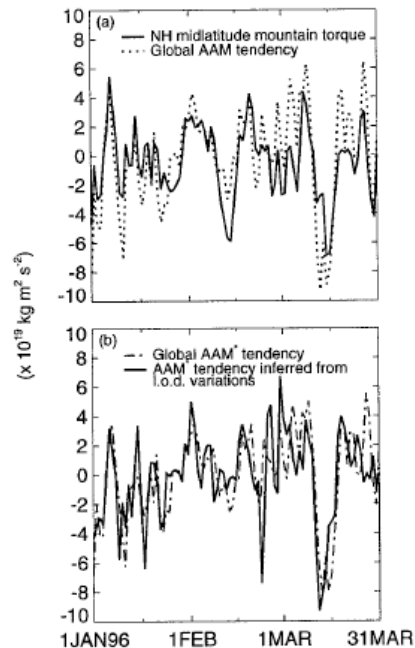
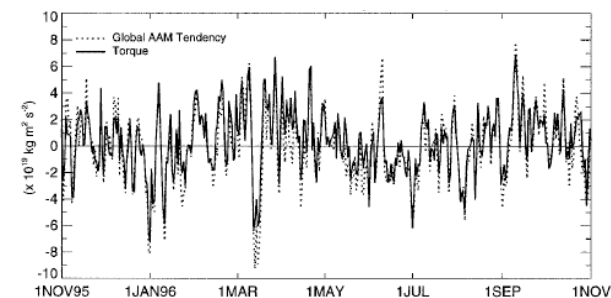


FIG. 8. Sea level pressure at 1200 UTC for (a) 8 March, (b) 15 March, and (c) 20 March 1996. Contours are every 6 hPa. Highs and lows discussed in the text are indicated by the letters H and L.

Surface Pressure Difference 15 March – 8 March

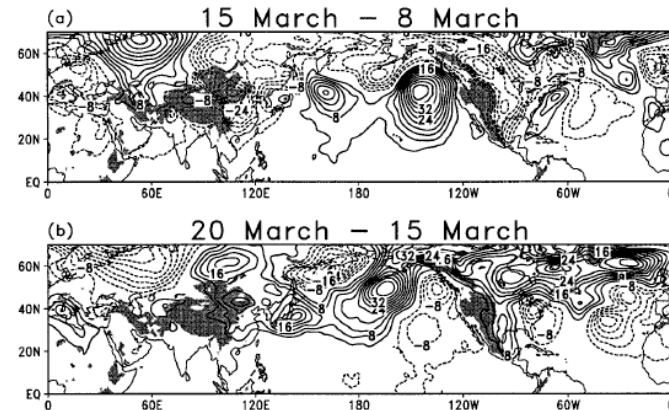
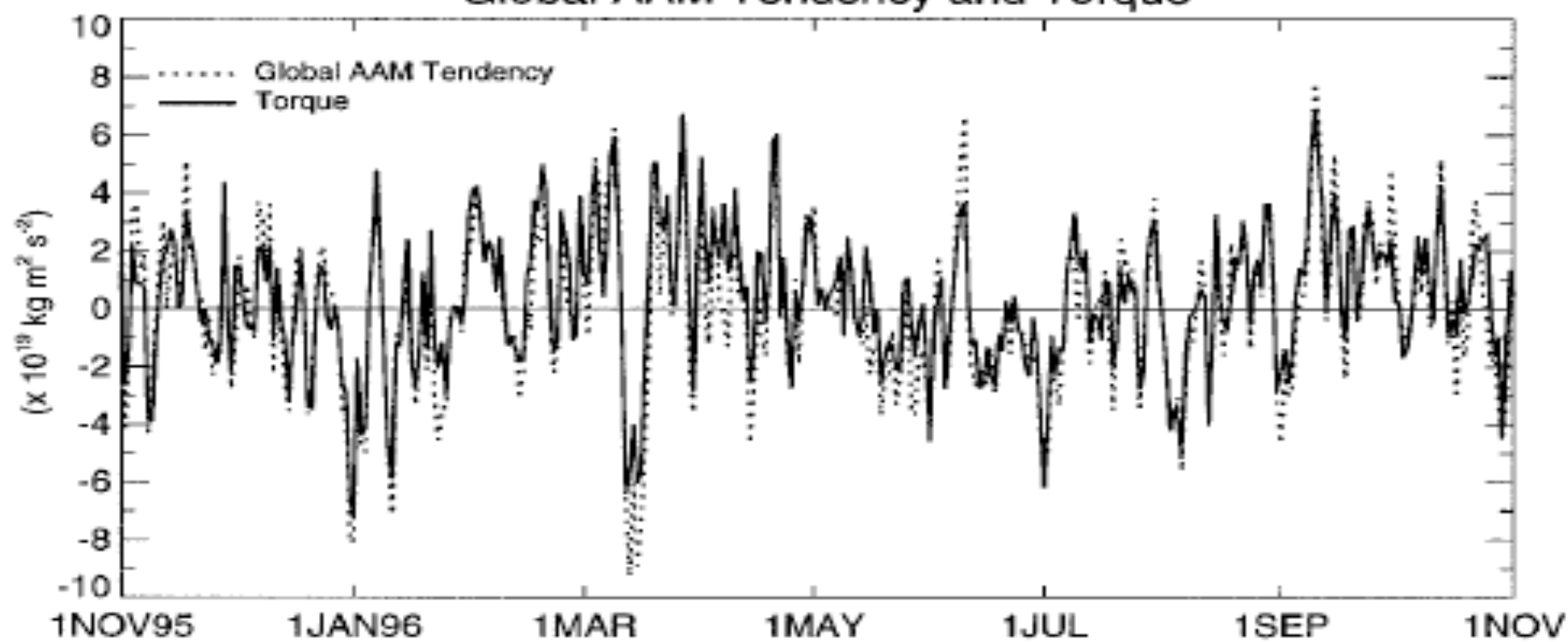


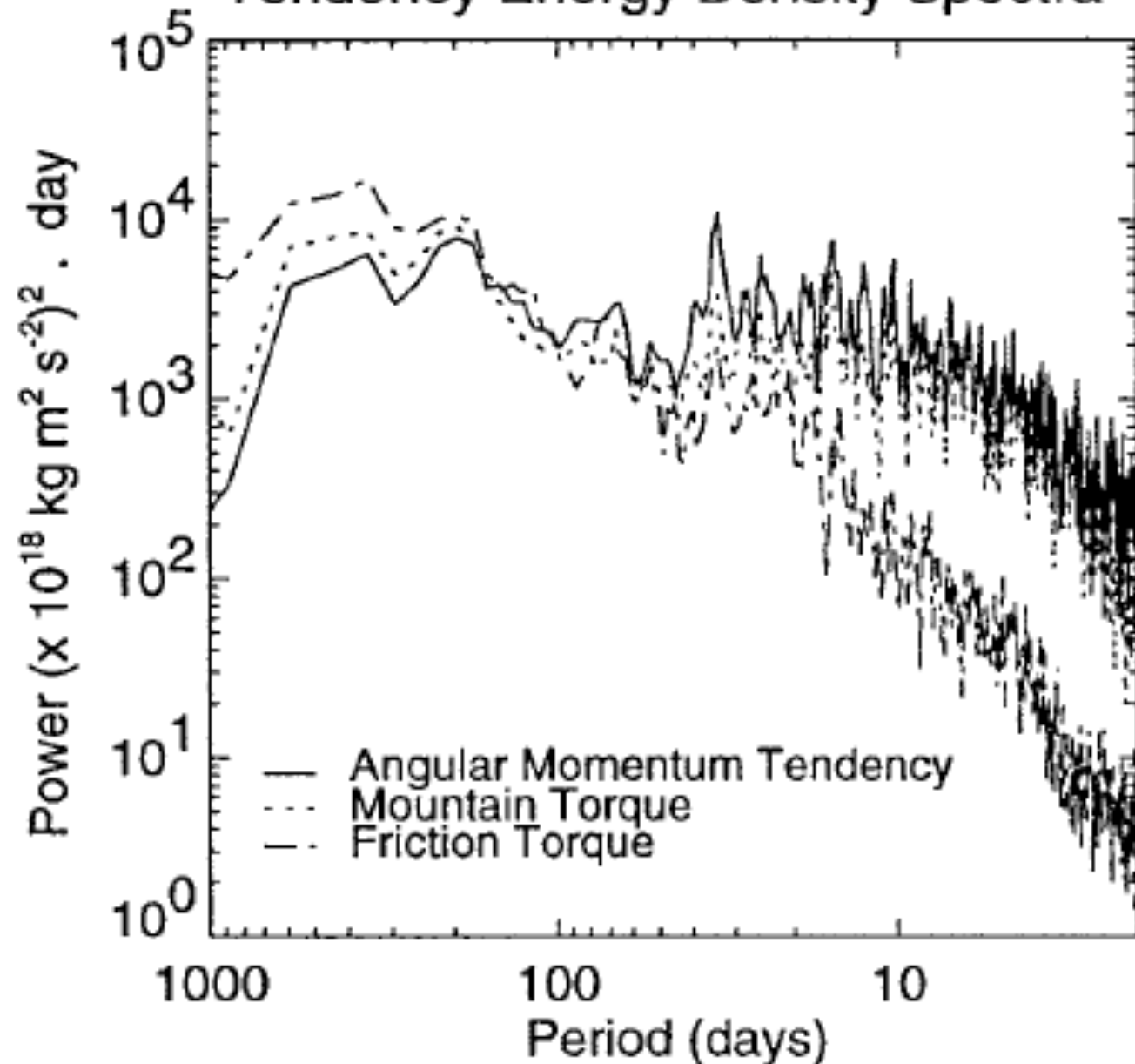
FIG. 9. (a) Difference in surface pressure between 1200 UTC 15 March and 1200 UTC 8 March 1996 (every 4 hPa contoured, zero line omitted). Positive contours are solid and negative dashed. (b) Same as (a) except between 1200 UTC 20 March and 1200 UTC 15 March. Shading indicates surface elevation of at least 1000 m.

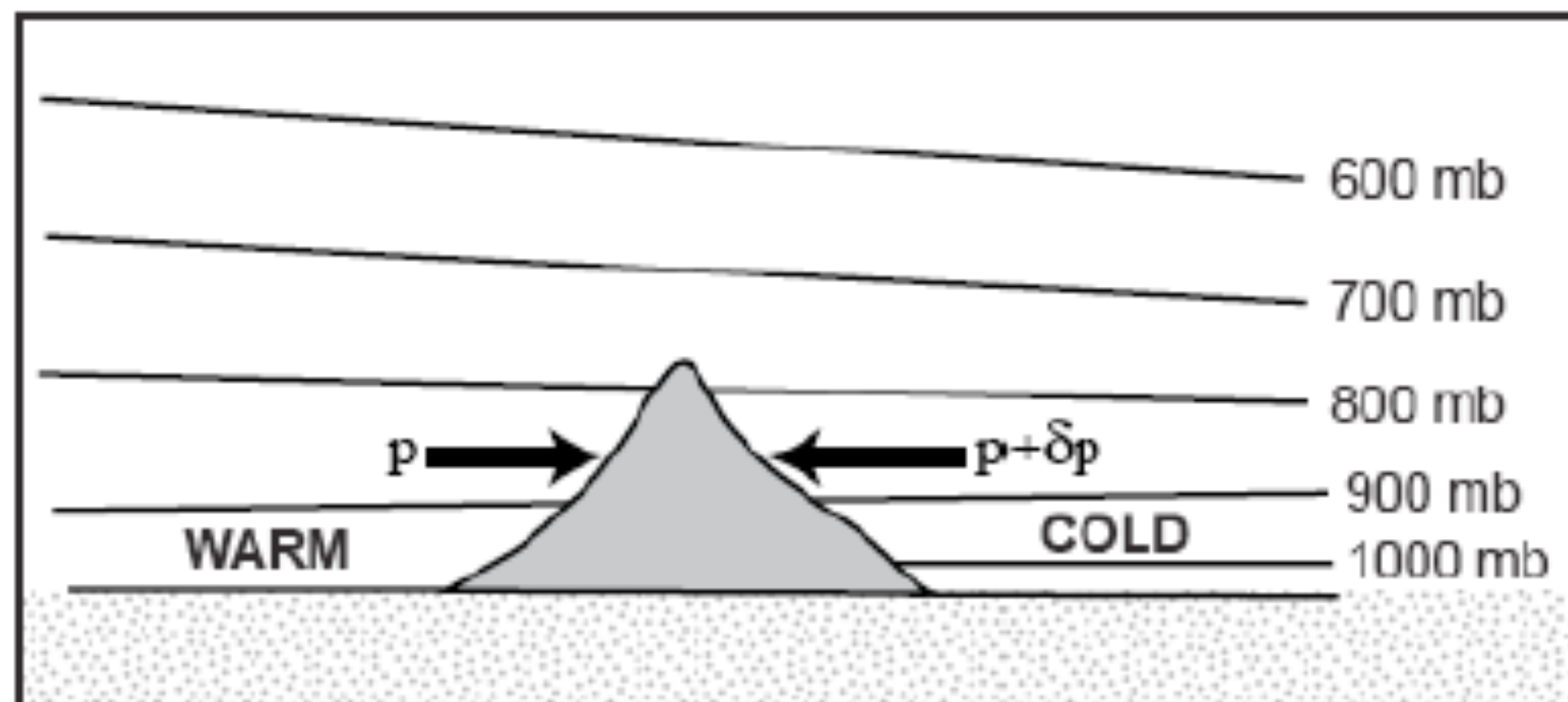
Global AAM Tendency and Torque



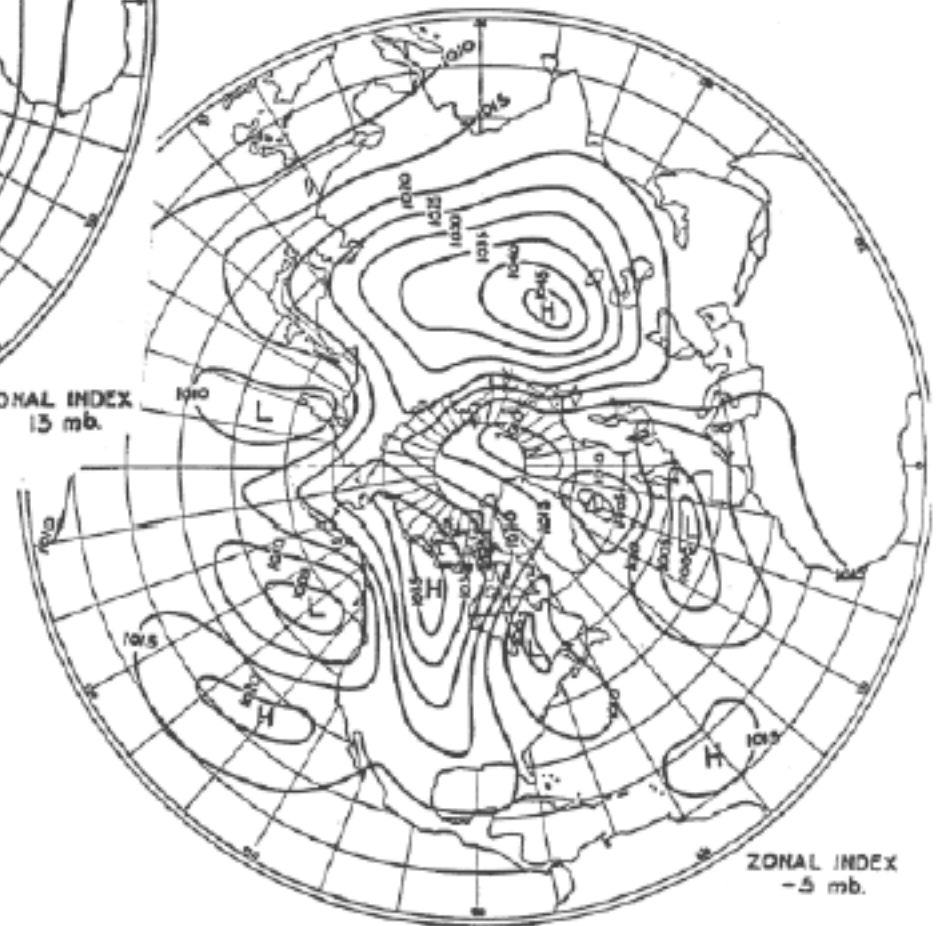
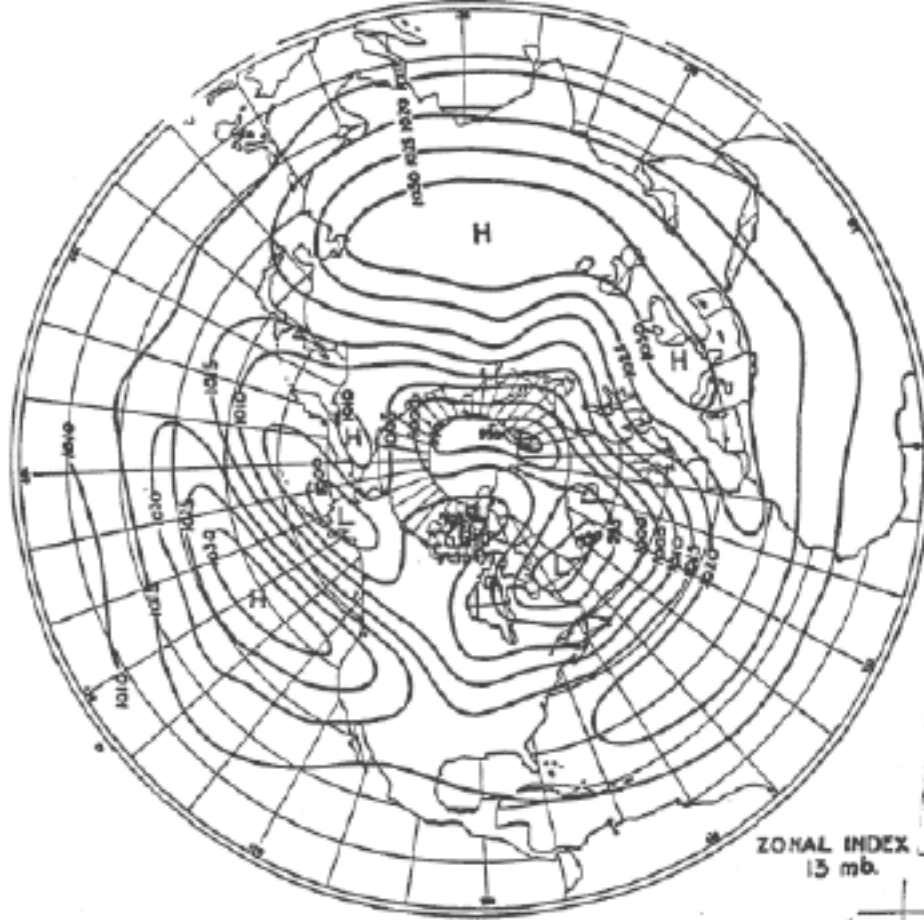
Global Torques and AAM

Tendency Energy Density Spectra

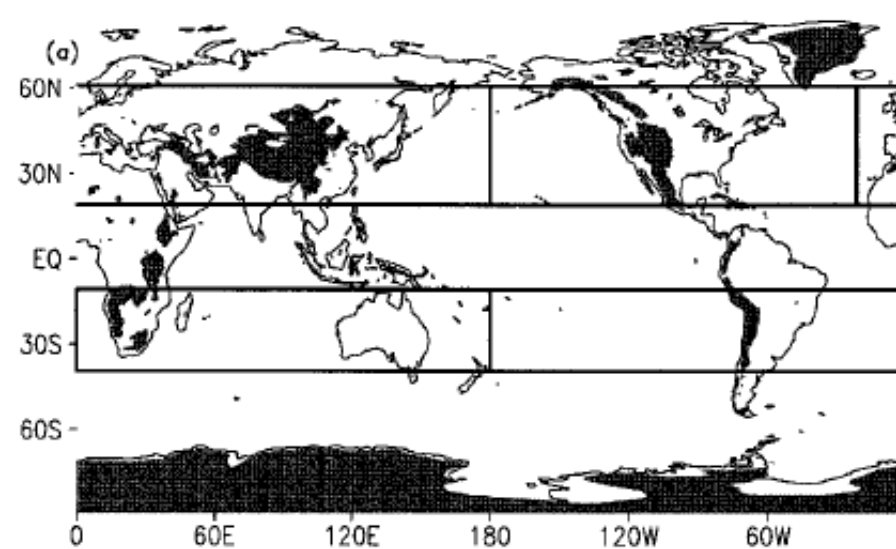
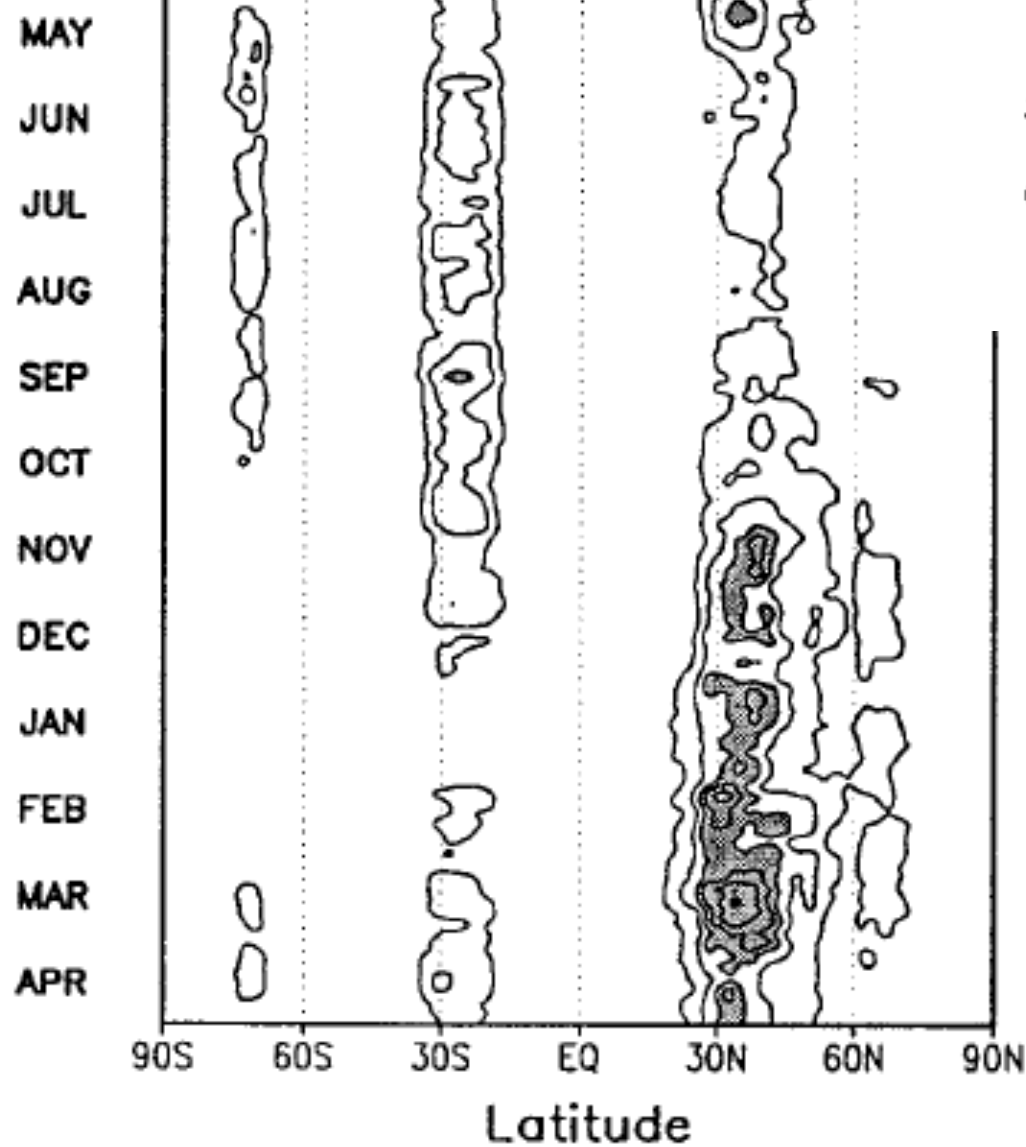




From Willett and Sanders

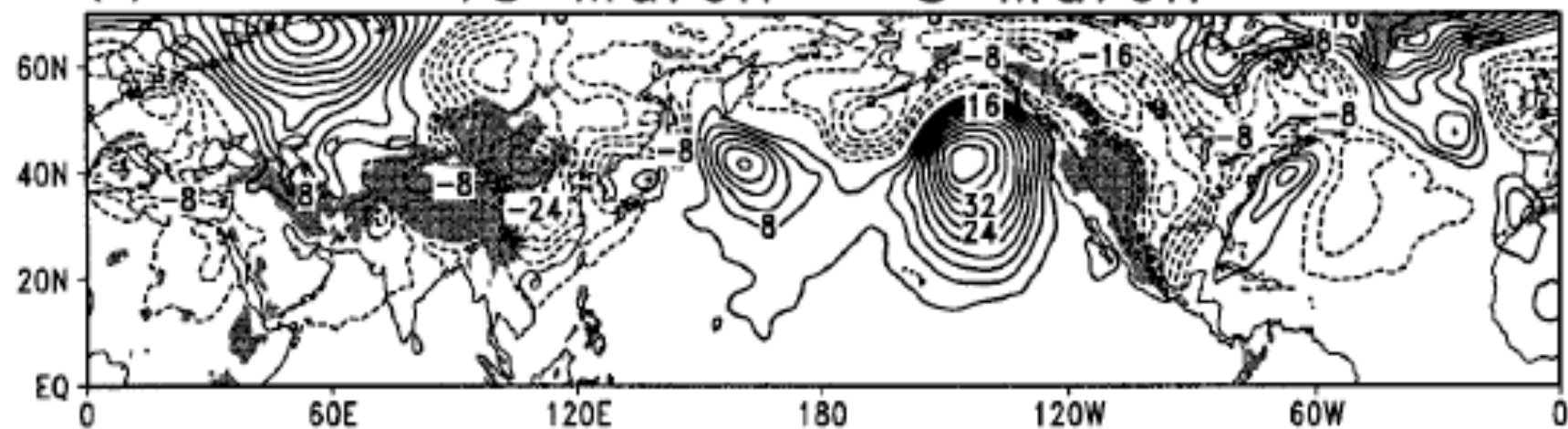


Mountain Torque Variance (14-day moving windows)



Surface Pressure Difference

(a) 15 March – 8 March



(b) 20 March – 15 March

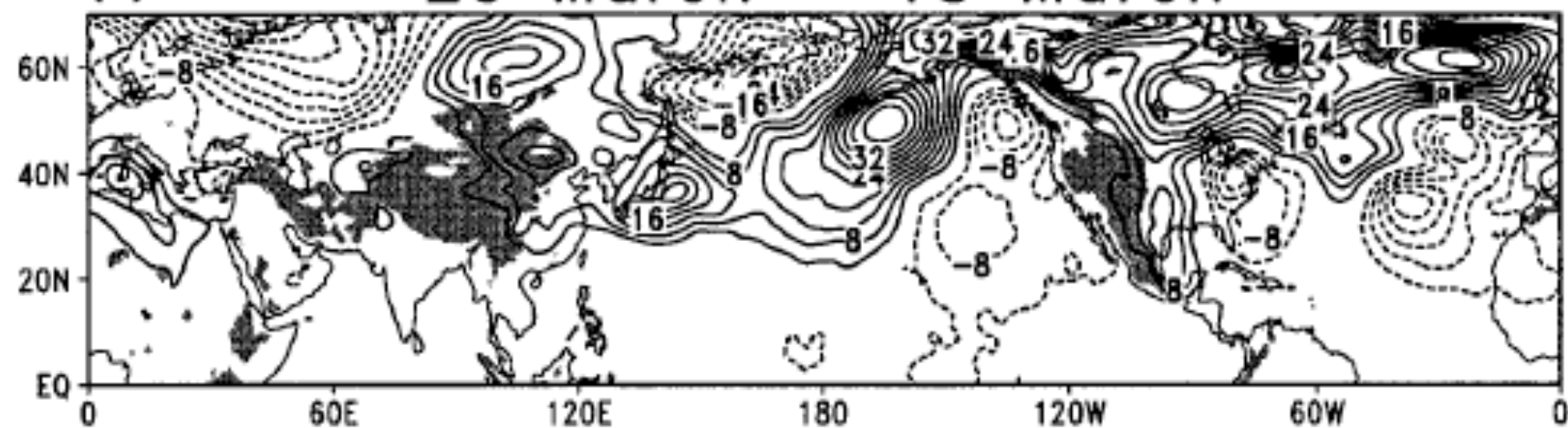
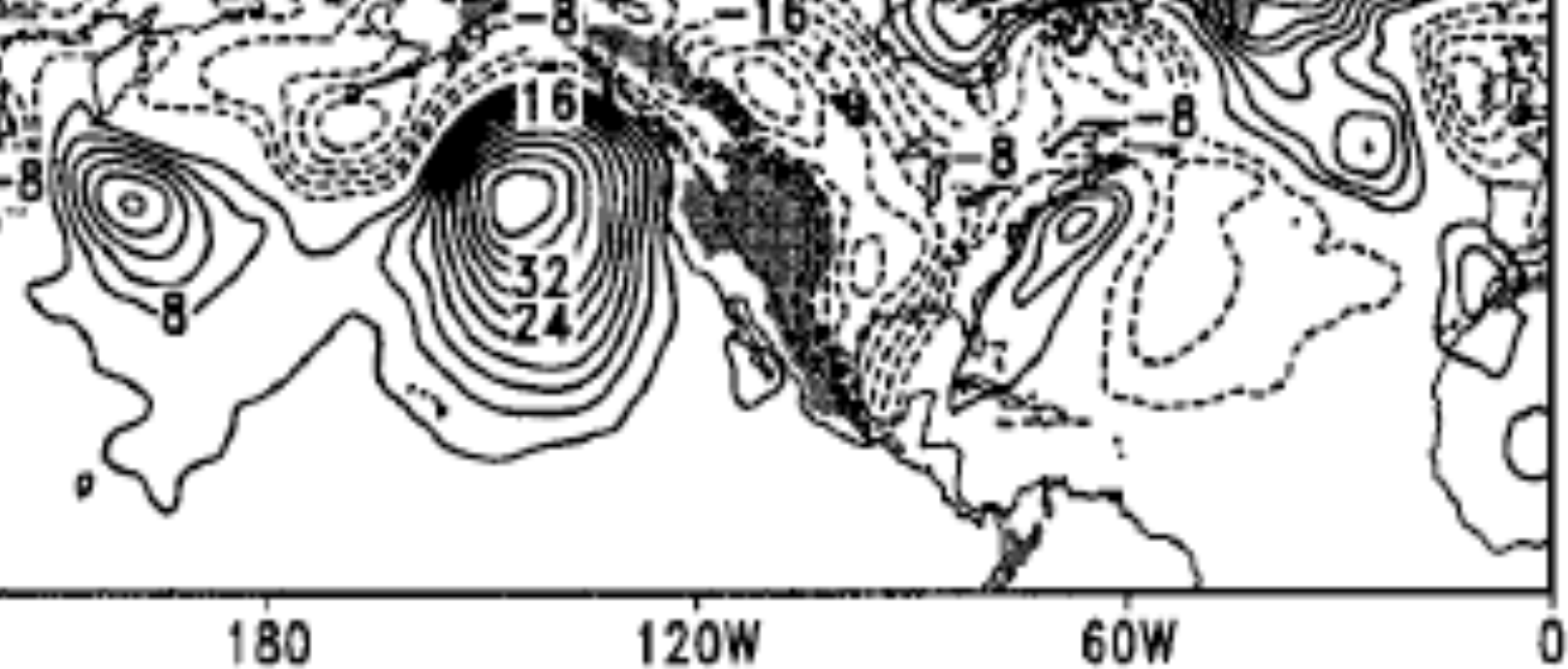
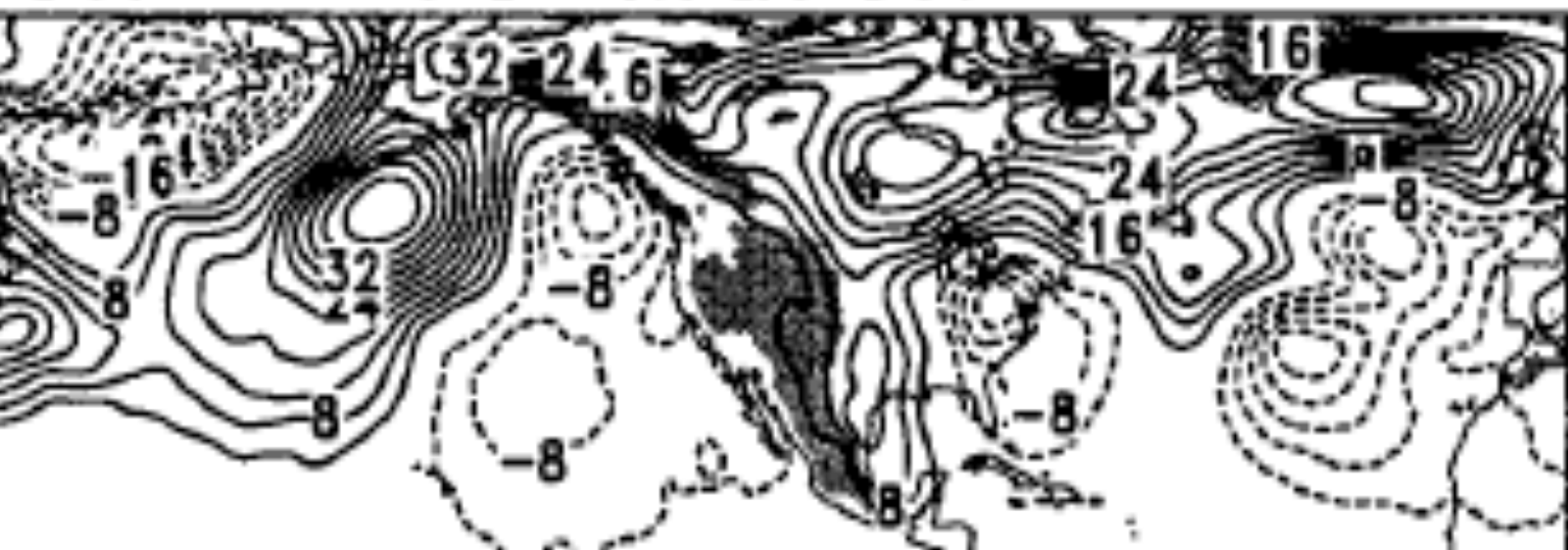


FIG. 9. (a) Difference in surface pressure between 1200 UTC 15 March and 1200 UTC 8 March 1996 (every 4 hPa contoured, zero line omitted). Positive contours are solid and negative dashed. (b) Same as (a) except between 1200 UTC 20 March and 1200 UTC 15 March. Shading indicates surface elevation of at least 1000 m.



ch - 15 March



For steady state:

Integrated over the globe

Net torque on atmosphere = 0

In the absence of mountains

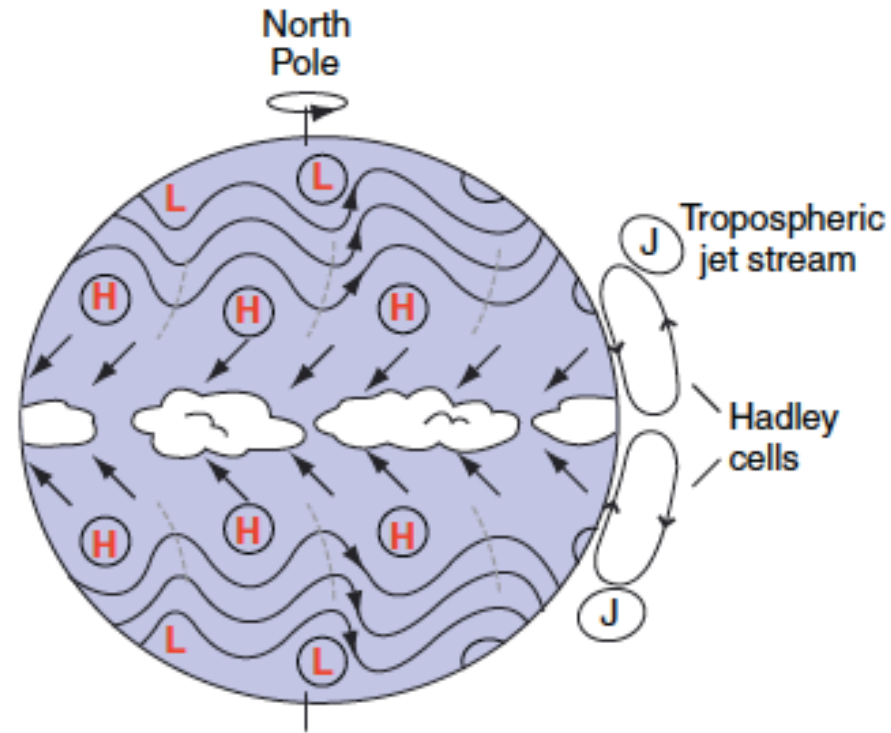
net frictional torque = 0

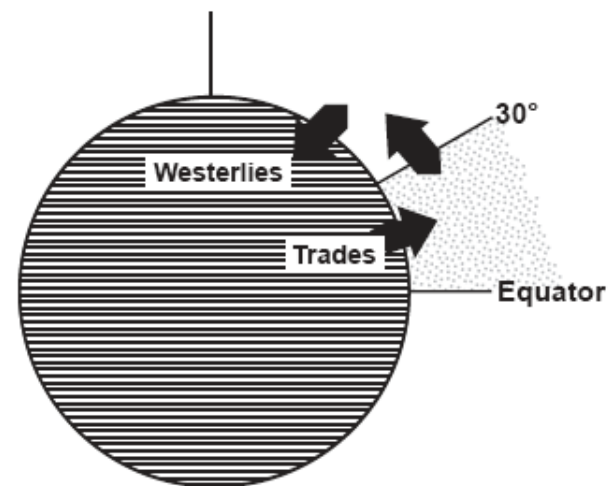
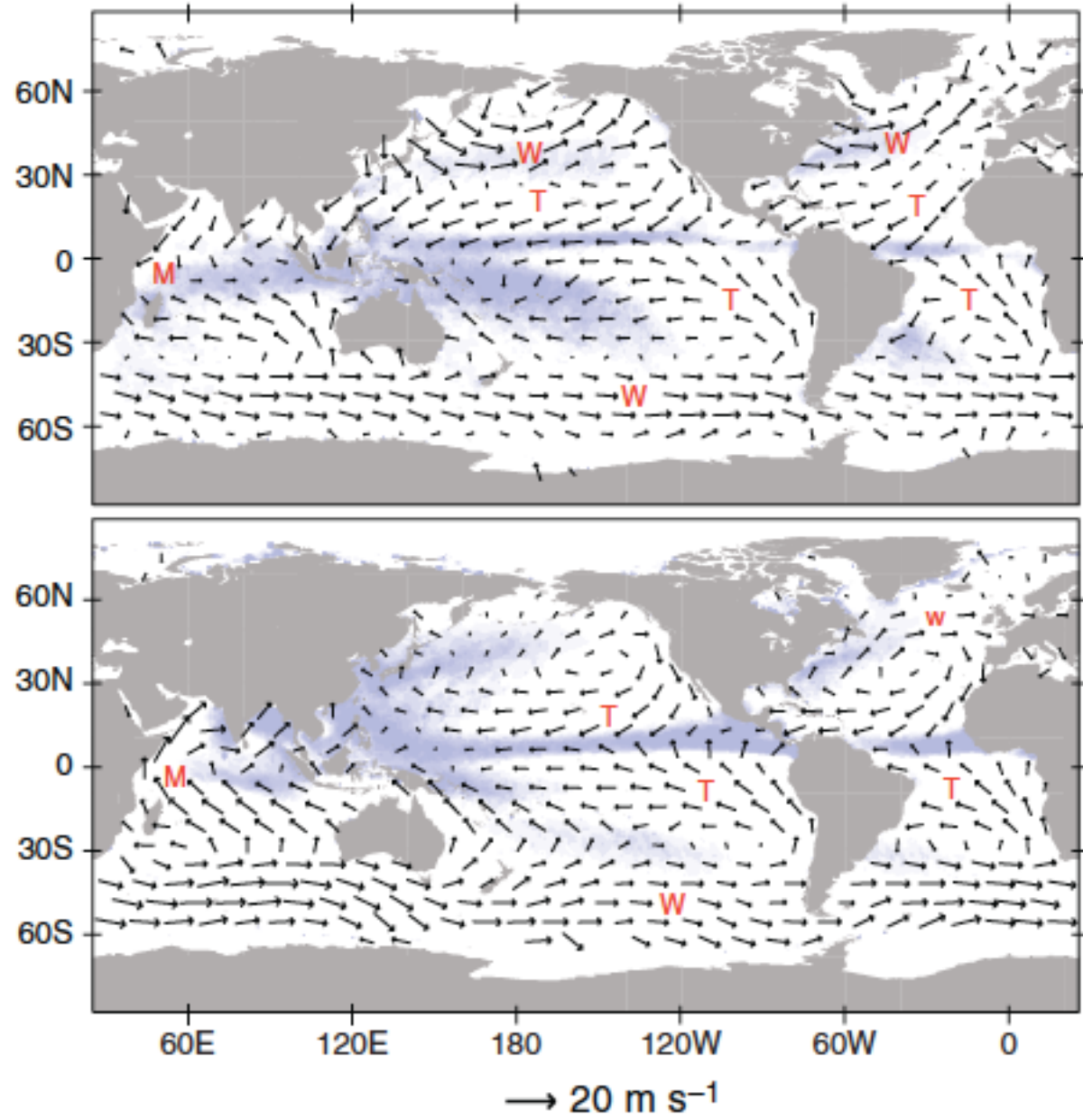
if the surface winds are nonzero, there must be regions of easterlies and westerlies

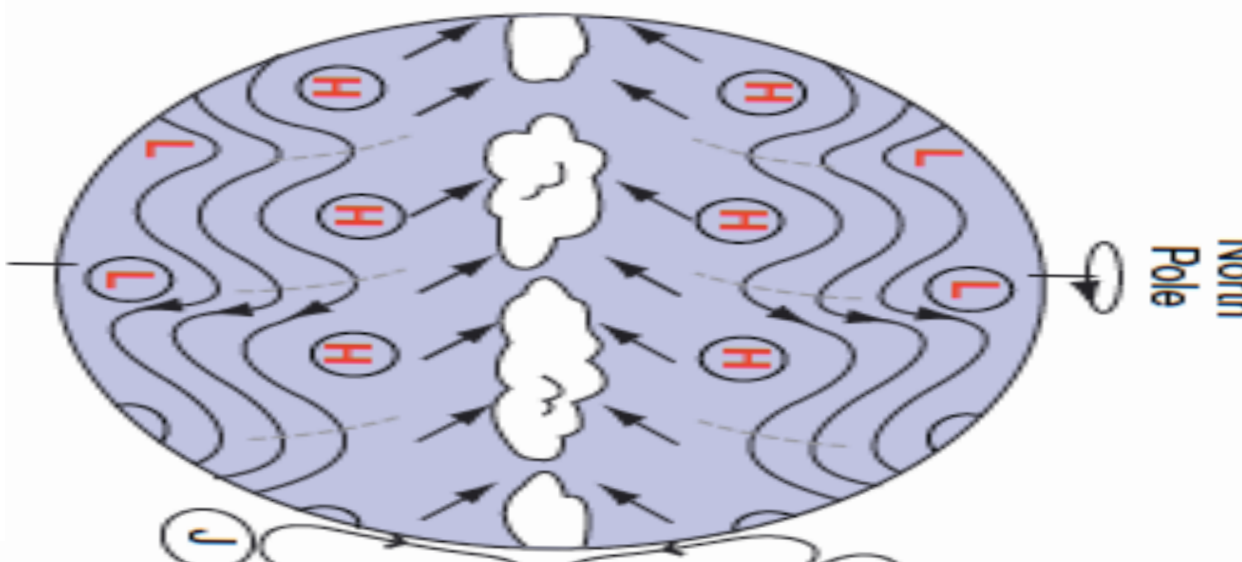
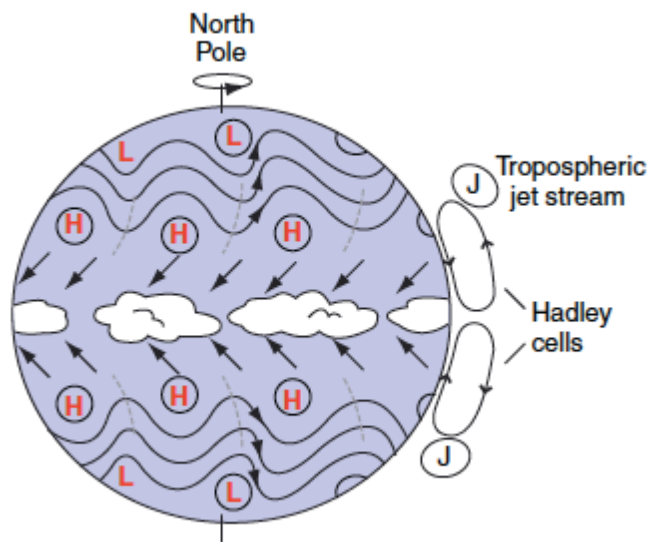
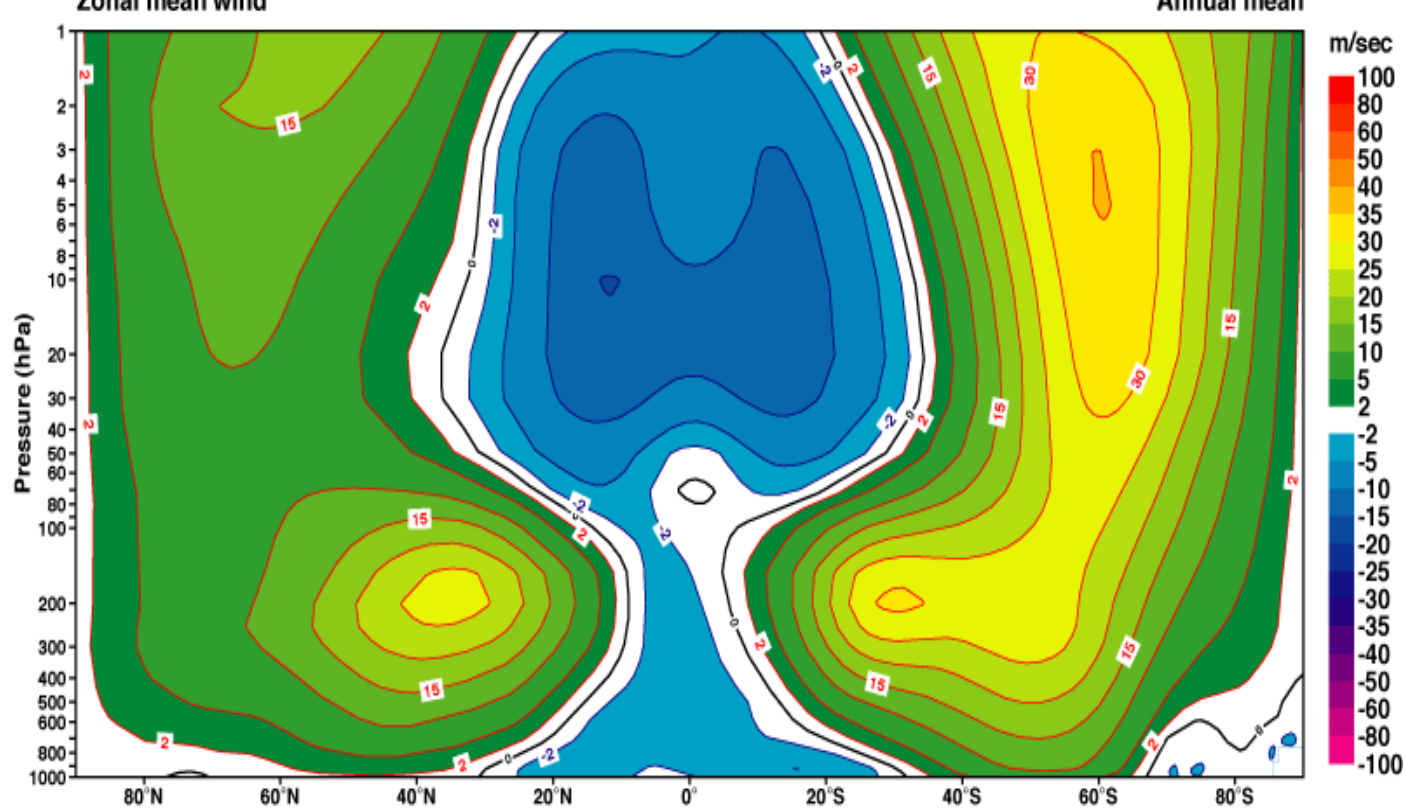
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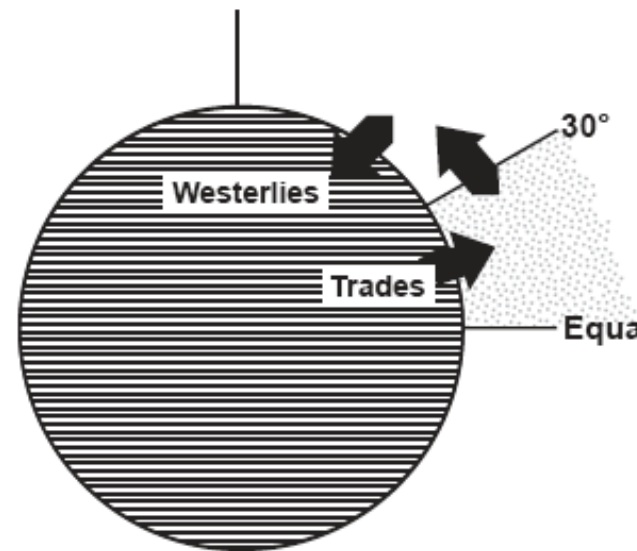
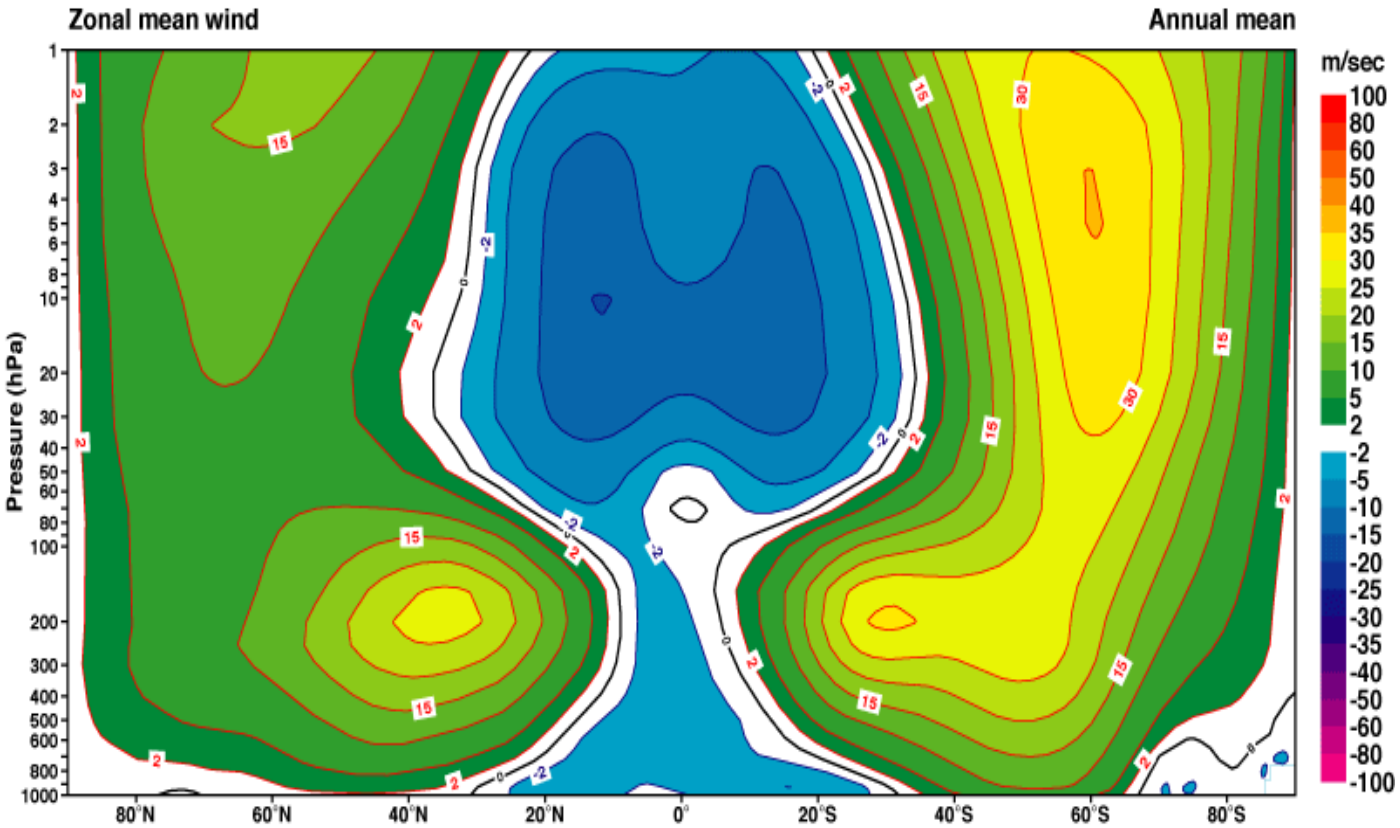
In the absence of mountains
net frictional torque = 0

if the surface winds are nonzero, there must be regions
of easterlies and westerlies



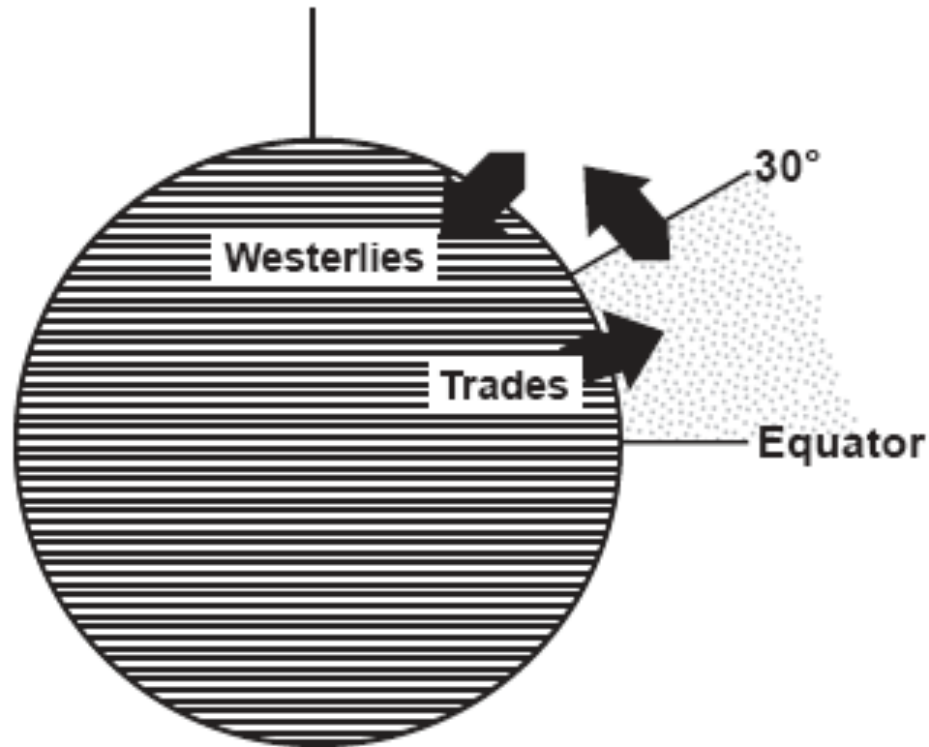






$$-2\pi R_E^3 \int_{eq}^{30^\circ} [\overline{\tau_x}] \cos^2 \phi d\phi = \frac{2\pi R_E \cos \phi}{g} \int_{30^\circ} [\overline{mv}] dp$$

Torque equatorward of 30° = Transport across 30°



Another balance requirement:

$$-2\pi R_E^3 \int_{eq}^{30^\circ} [\overline{\tau_x}] \cos^2 \phi d\phi = \frac{2\pi R_E \cos \phi}{g} \int_{30^\circ} [\overline{mv}] dp$$

Torque equatorward of 30° = Transport across 30°

It follows that there must be a poleward flux of AAM across 30° latitude in both hemispheres

$$\frac{2\pi R_E \cos \phi}{g} \int_{30^\circ} [\overline{mv}] dp = \frac{2\pi \Omega R_E^3 \cos^3 \phi}{g} \int_{30^\circ} [\overline{v}] dx dp + \frac{2\pi R_E^2 \cos^2 \phi}{g} \int_{30^\circ} [\overline{uv}] dx dp$$

Total

M_Ω

M_r

*requires
poleward
mass flux*

Decomposition of M_r

$$[\overline{uv}] = [\overline{u}] [\overline{v}] + \overline{[u]'} [v] + [\overline{u}^* \overline{v}^*] + \overline{[u^{*'} v^{*'}]}$$

Steady
MMC

transient
MMC

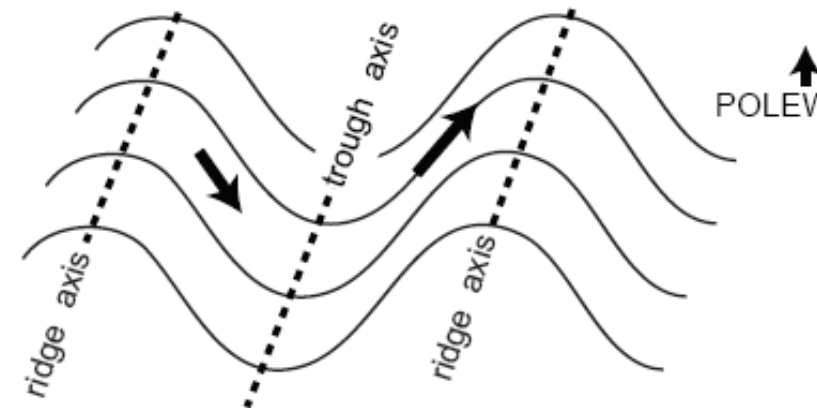
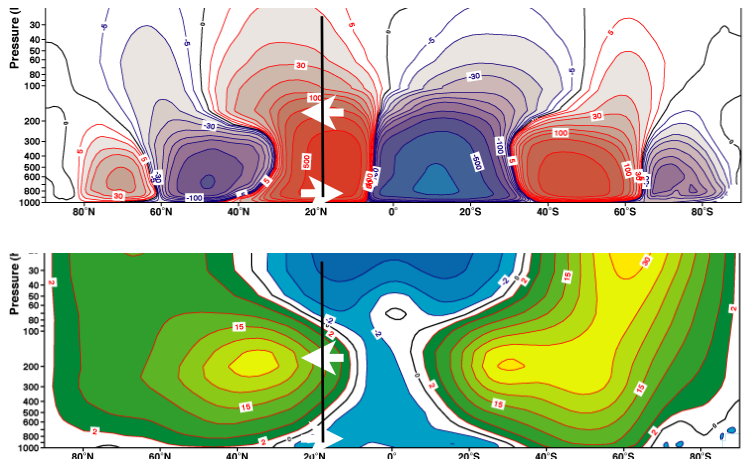
steady
eddy

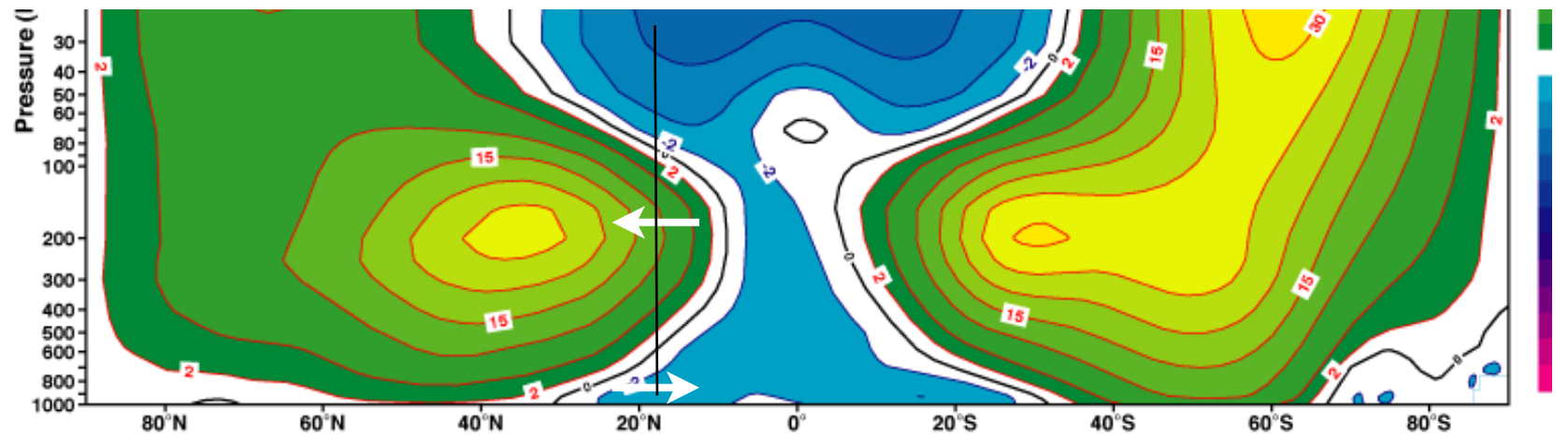
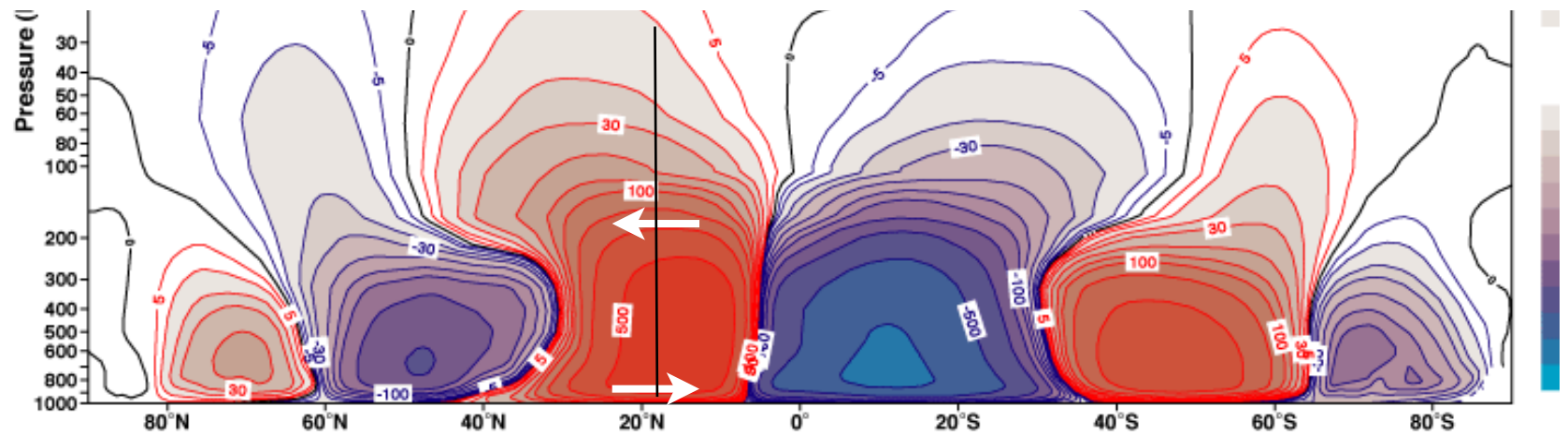
transient
eddy

a.k.a.
stationary wave

Decomposition of M_r

$$[\overline{uv}] = [\overline{u}] [\overline{v}] + \overline{[u]'} [v] + [\overline{u^*} \overline{v^*}] + \left[\overline{u^{*'} v^{*'}} \right]$$

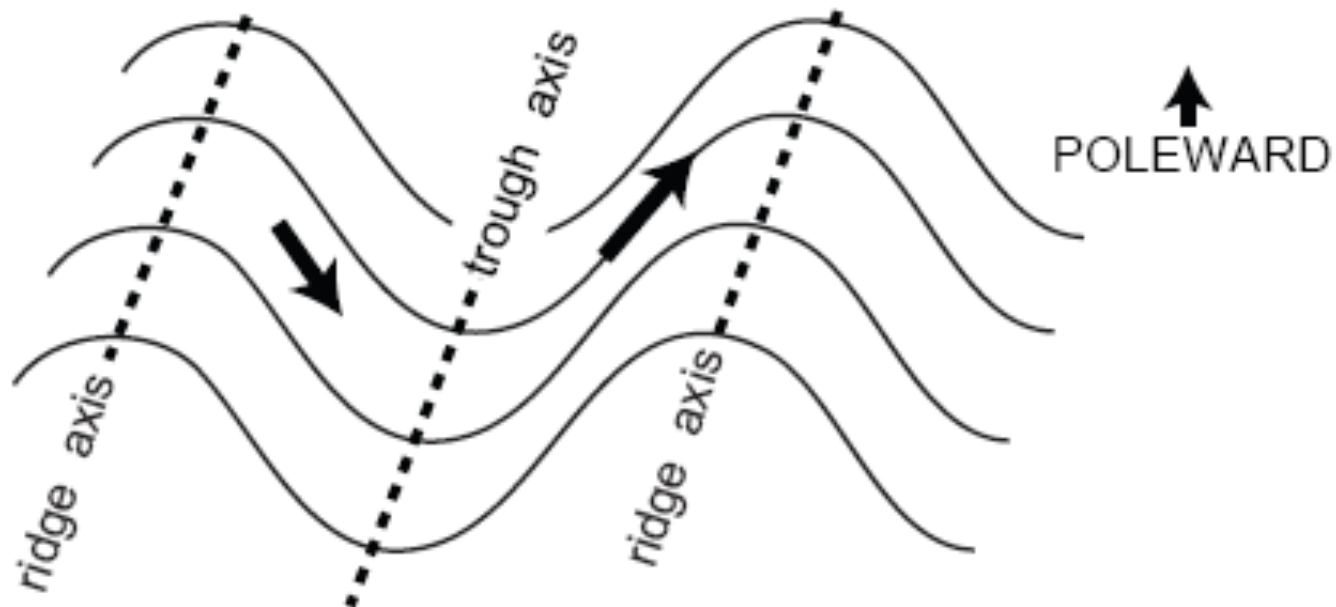




$$\int_0^{p_0} [u][v] dp$$

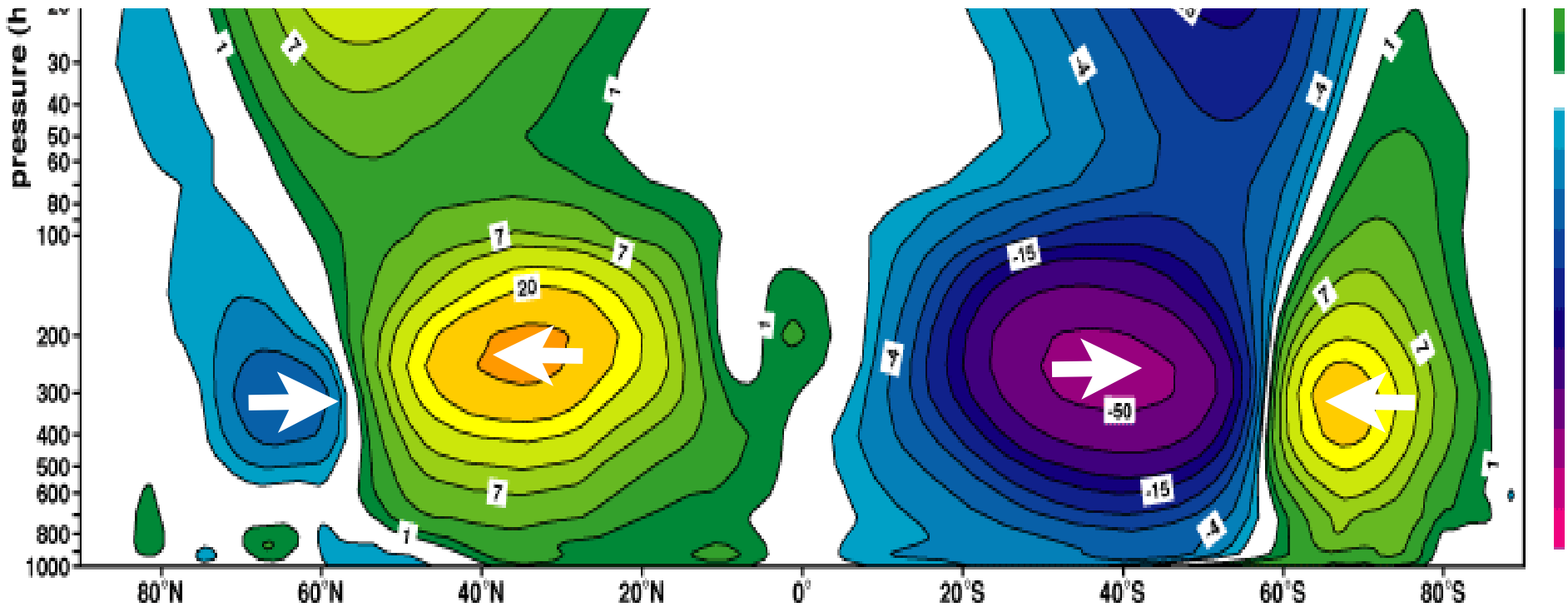
Product Moment Formula

$$\left[u * v * \right] = r(u, v) \times \sigma_u \times \sigma_v$$

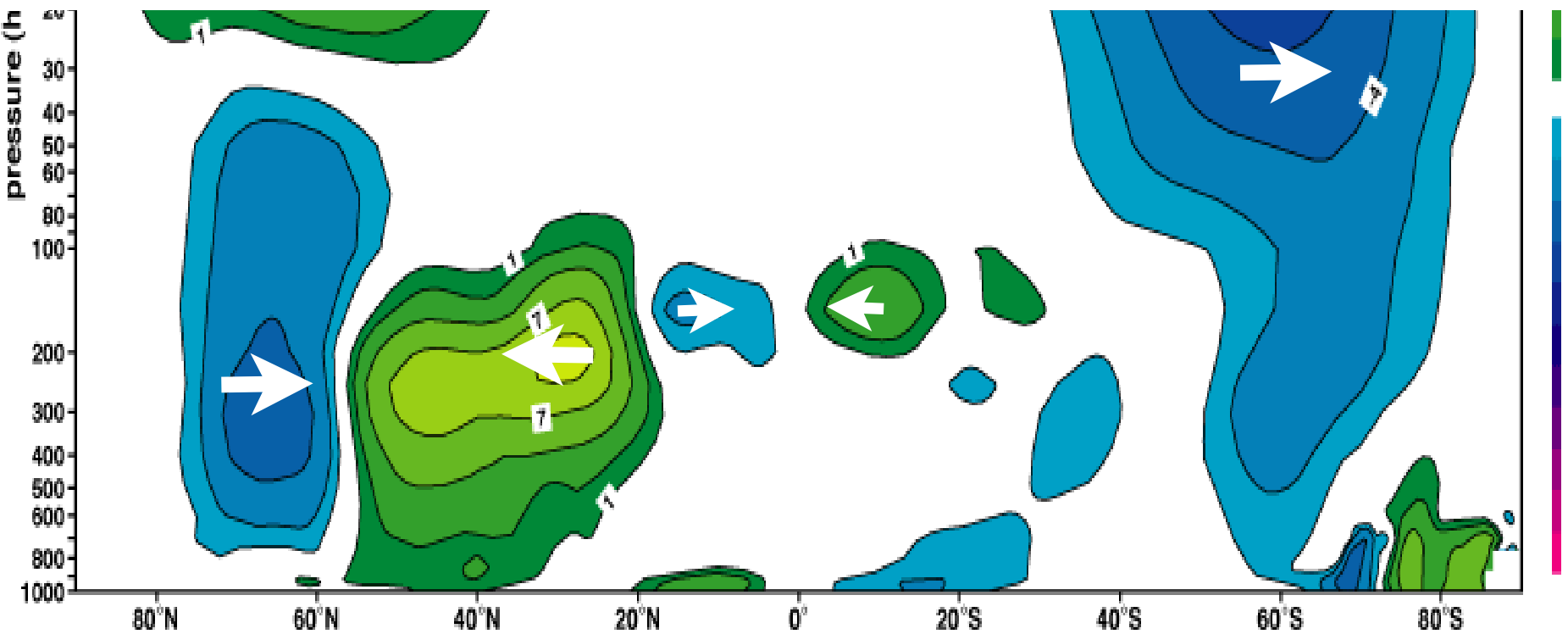


$$[u * v *]$$

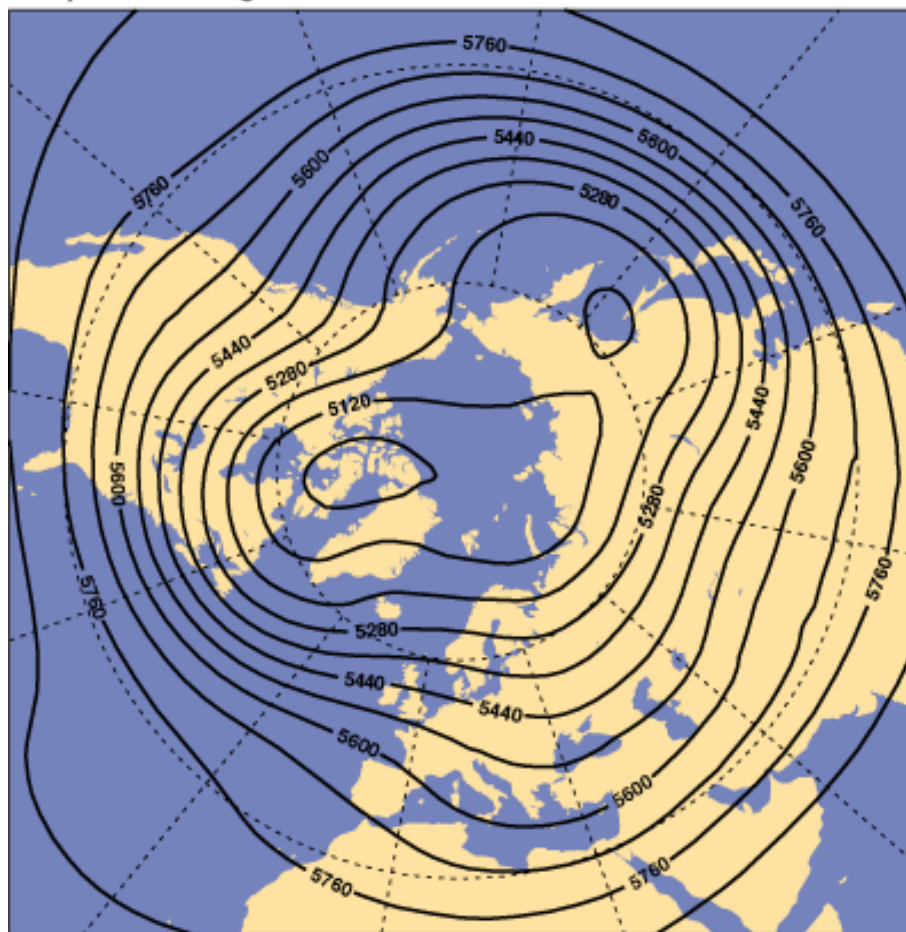
transient eddy component annual mean



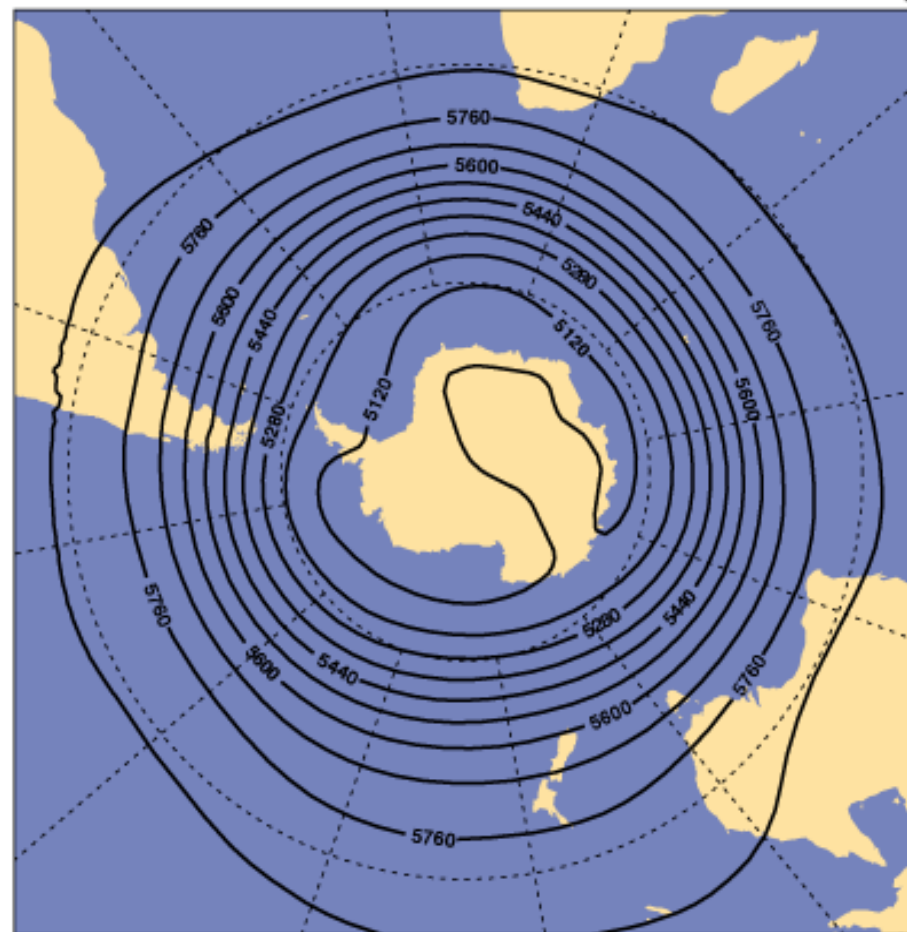
stationary wave component annual mean

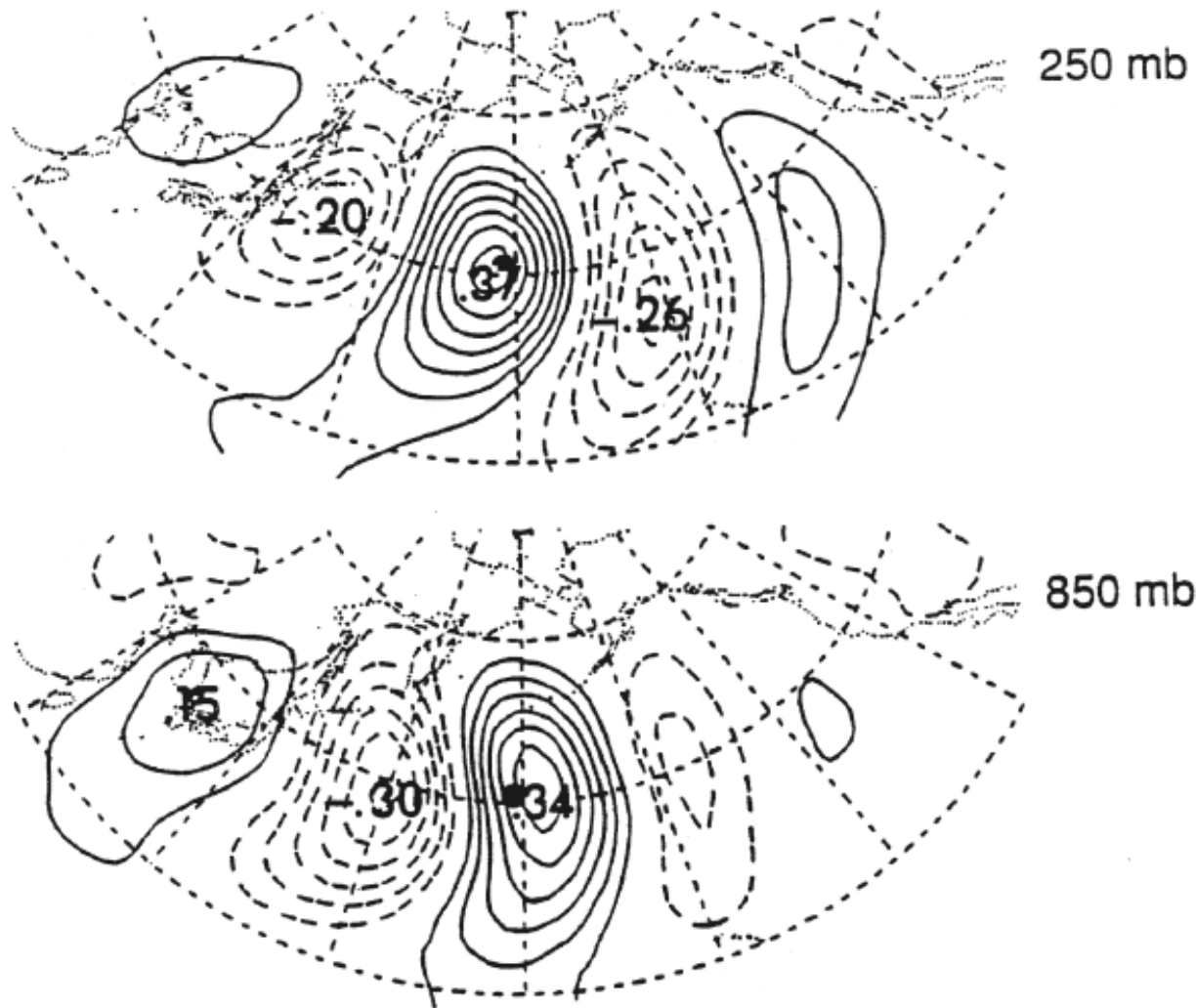


Geopotential height at 500 hPa



December-February





one-point correlation maps; 500 hPa height; highpass filtered

Conclusions

In accordance with the balance requirements, there is a strong poleward flux of angular momentum across 30° latitude

The flux is greater during winter when the surface westerlies are stronger

The poleward flux across 30° is accomplished **exclusively by the eddies**

Transient eddies and stationary waves both contribute

Nearly all the flux occurs around the jet stream level (above 500 hPa)

Balance requirement for an upward flux of M equatorward of 30° and a downward flux poleward of 30°

Vertical transport of angular momentum

$$\frac{2\pi\Omega R_E^3}{g} \int -[\bar{\omega}] \cos^3 \phi dy + \frac{2\pi R_E^2}{g} \int -[\overline{u\omega}] \cos^2 \phi dy$$

$$M_{\Omega}$$

$$M_r$$

$$\frac{2\pi R_E^2}{g} \int -[\bar{\omega}] (\Omega R_E \cos \phi + [\bar{u}]) \cos^2 \phi dy + \frac{2\pi R_E^2}{g} \int [\overline{u^* \omega^*}] \cos^2 \phi dp$$

MMC term

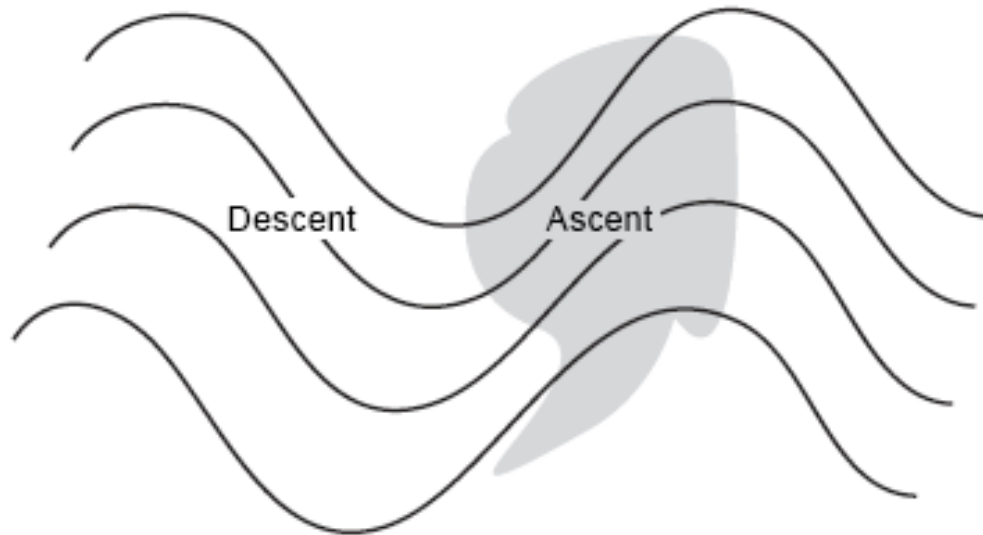
eddy term

The eddies are not the answer

Scaling arguments: extratropical eddy fluxes are too small by a factor of Ro

Tropical eddy fluxes are almost nonexistent

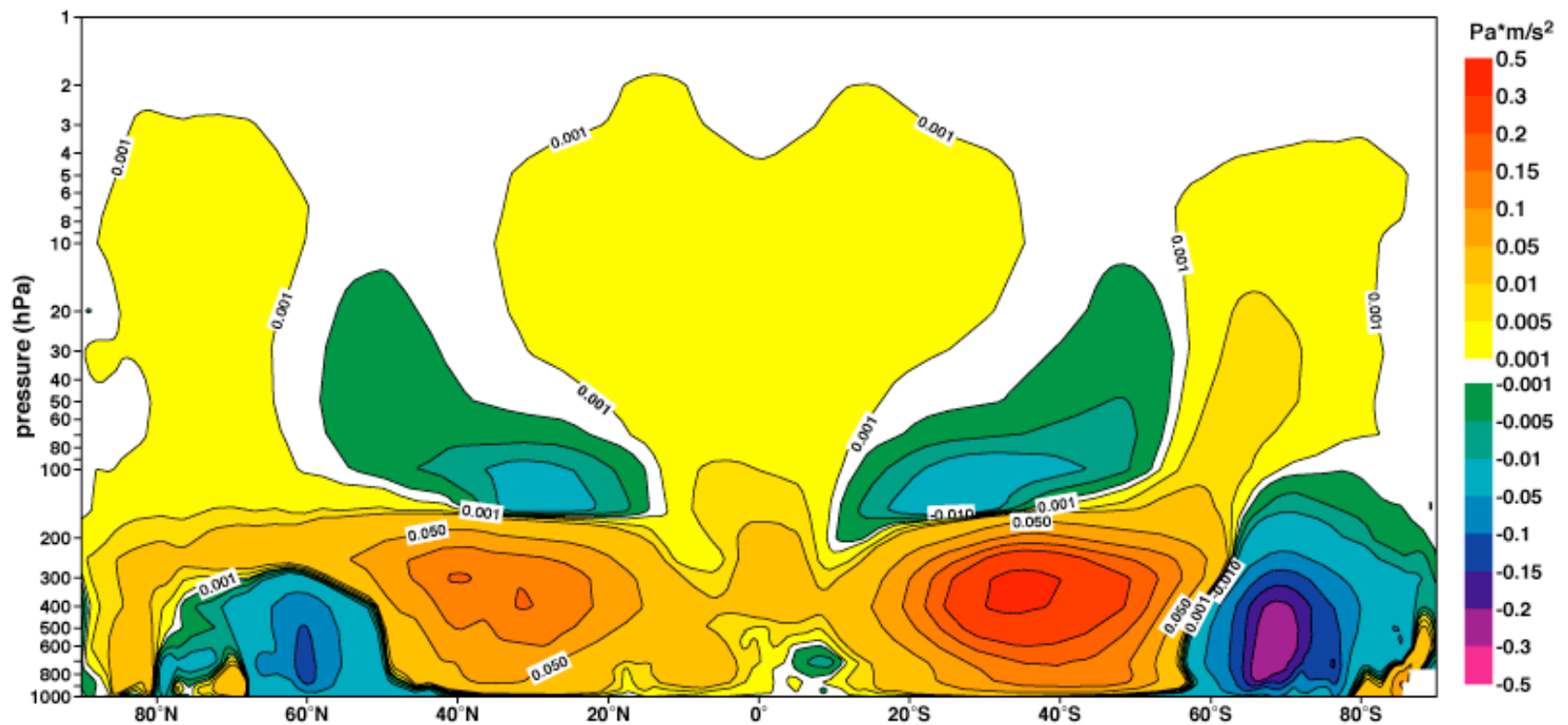
Extratropical eddy fluxes are upward; not downward



So it must be the MMC

Transient upward eddy flux of westerly wind

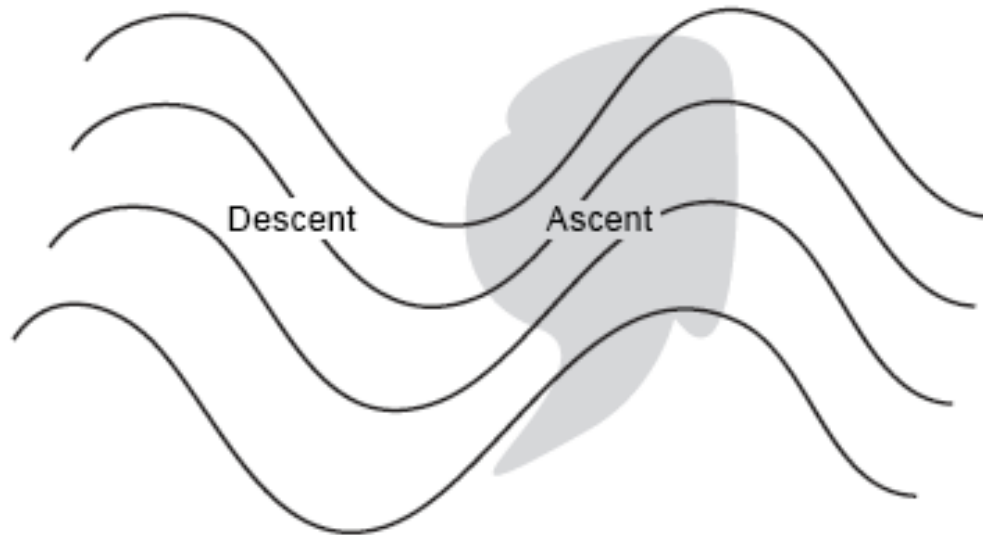
Annual mean



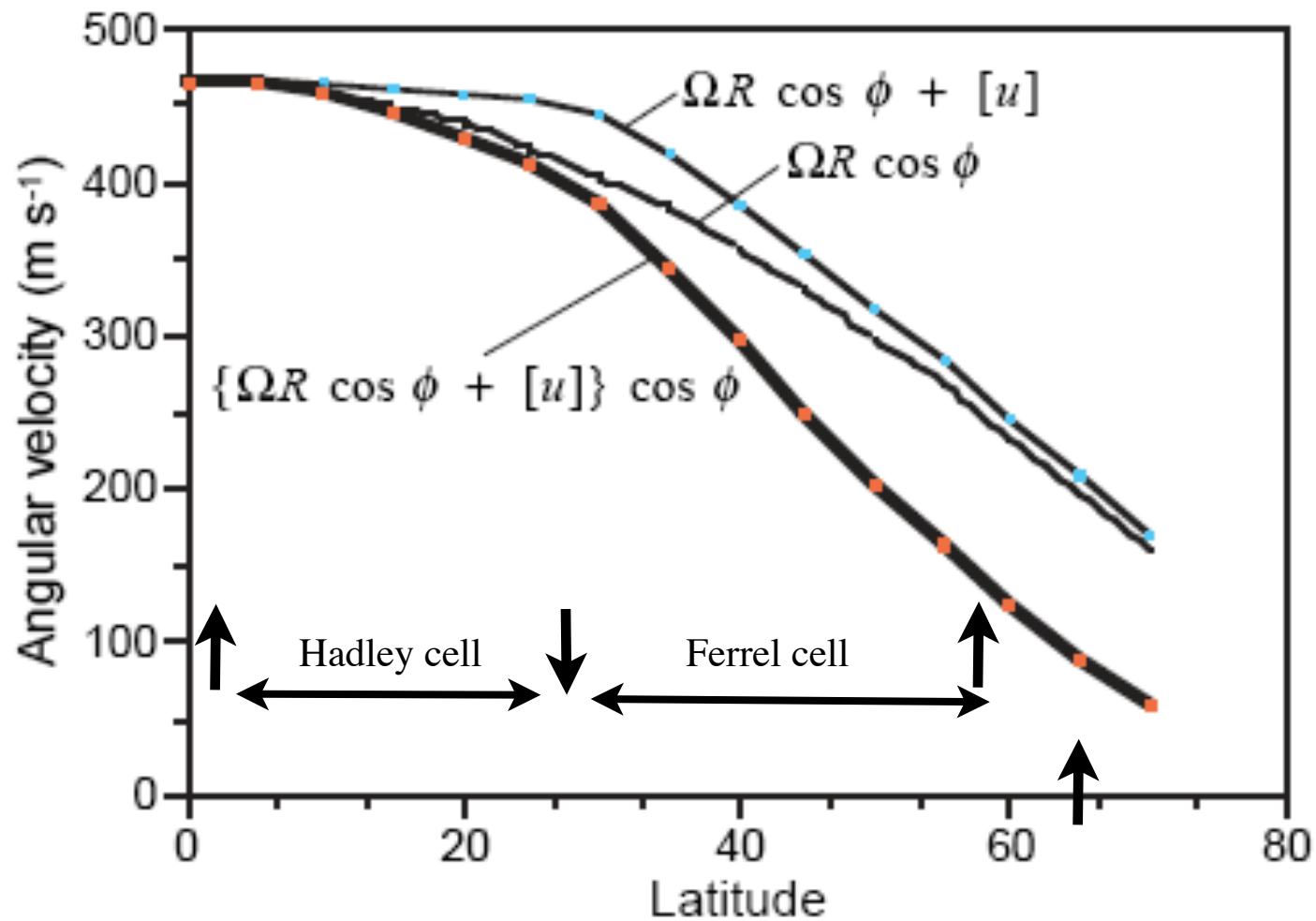
The eddies are not the answer

Scaling arguments: extratropical eddy fluxes are too small by a factor of Ro

Extratropical eddy fluxes are upward; not downward



So it must be the MMC



$$\frac{2\pi R_E^2}{g} \int -[\bar{\omega}] (\Omega R_E \cos \phi + [u]) \cos^2 \phi dy$$

“Spin down” of the circulation in a teacup

The zonally averaged equation of motion

Spherical geometry

$$\frac{\partial [u]}{\partial t} = [v] \left(f - \frac{1}{\cos \phi} \frac{\partial}{\partial y} [u] \cos \phi \right) - [\omega] \frac{\partial [u]}{\partial p} - \frac{1}{\cos^2 \phi} \frac{\partial}{\partial \phi} [u^* v^*] \cos^2 \phi - \frac{\partial}{\partial p} [u^* \omega^*] + F_x$$

Cartesian geometry

$$\frac{\partial [u]}{\partial t} = [v] \left(f - \frac{\partial [u]}{\partial y} \right) - [\omega] \frac{\partial [u]}{\partial p} - \frac{\partial}{\partial y} [u^* v^*] - \frac{\partial}{\partial p} [u^* \omega^*] + F_x$$

Neglecting vertical advection by MMC; using G to represent eddies

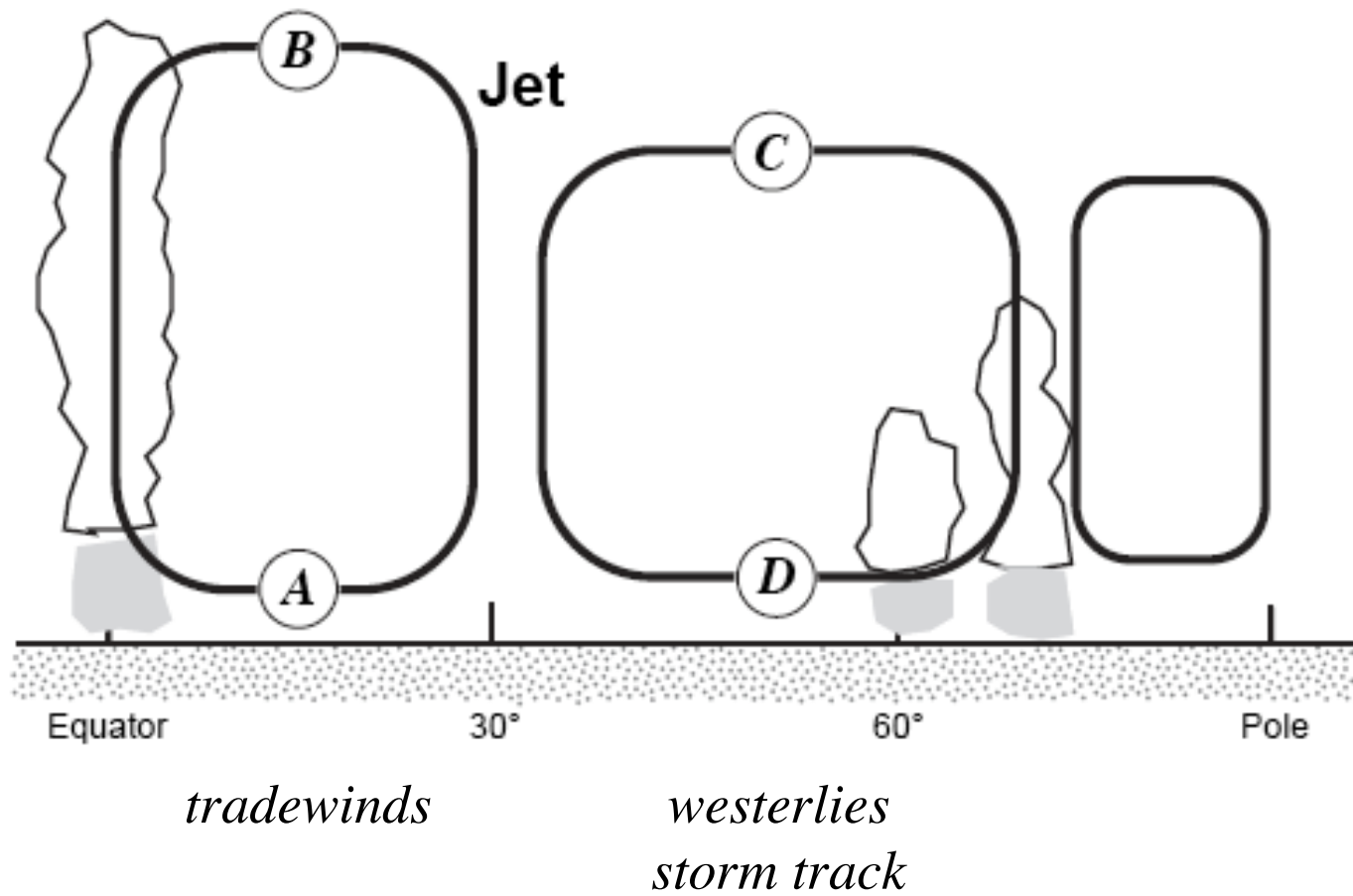
$$\frac{\partial [u]}{\partial t} = [v] \left(f - \frac{\partial [u]}{\partial y} \right) + G + F_x$$

$$\frac{\partial [u]}{\partial t} = [v] \left(f - \frac{\partial [u]}{\partial y} \right) + G + F_x$$

MMC *dynamic stability* eddy frictional
 $\propto dM/dy$ forcing drag

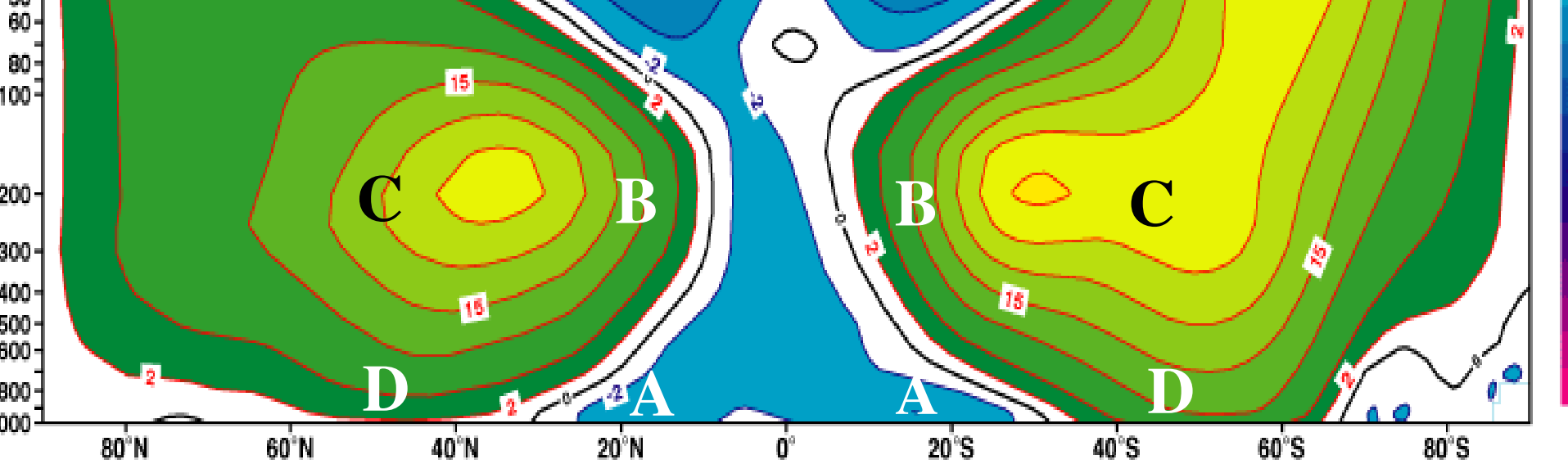
For long term “balance requirement” $d/dt = 0$

$$[v] = \frac{G + F_x}{f - \partial[u] / \partial y}$$



at A and D $G = 0$

at B and C $F = 0$



at **B** $d[u]/dy \sim 30 \text{ m s}^{-1}$ over 2000 km $\sim 1.5 \times 10^{-5} \text{ s}^{-1}$

$$f \sim 4 \times 10^{-5} \text{ s}^{-1}$$

$$f - d[u]/dy \sim 2.5 \times 10^{-5} \text{ s}^{-1}$$

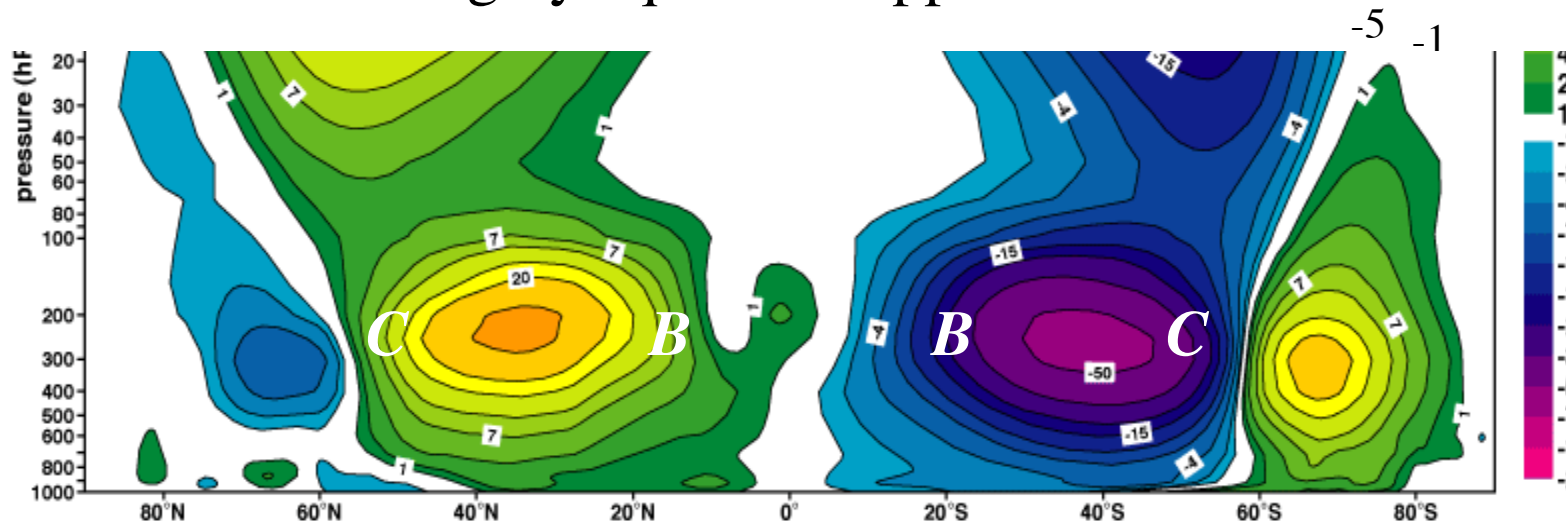
at **C** $f - d[u]/dy \sim 10 \times 10^{-5} \text{ s}^{-1}$ $4 \times$ larger than at B

Recalling that

$$[v] = \frac{G + F_x}{f - \partial[u] / \partial y}$$

and $F_x = 0$ at **B** and **C**

and G is roughly equal and opposite at **B** and **C**



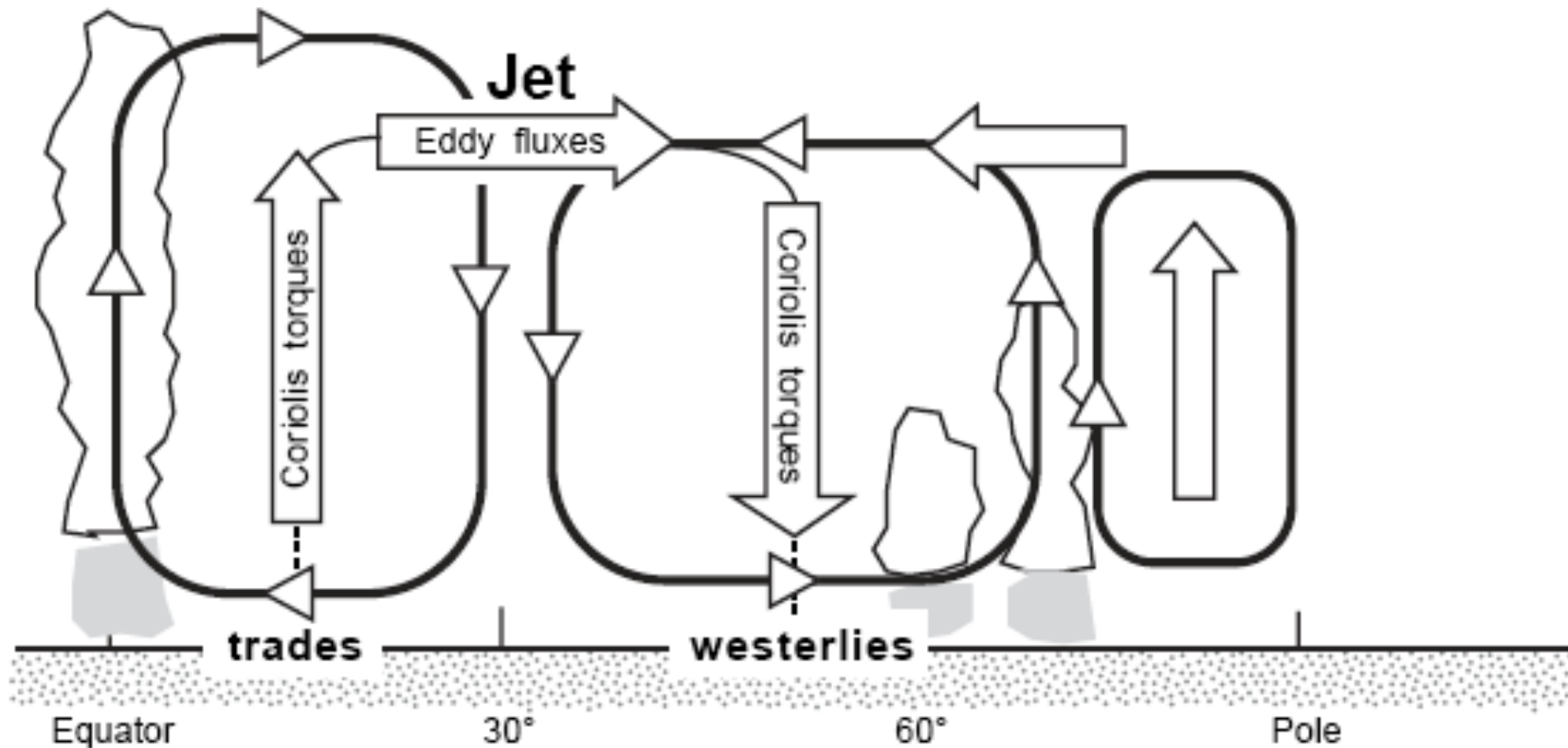
Recalling that

$$[v] = \frac{G + F_x}{f - \partial[u] / \partial y}$$

and $F_x = 0$ at ***B*** and ***C***

and G is roughly comparable at ***B*** and ***C***

**it follows that $[v]$ is $\sim 4 \times$ stronger at *B* than at *C*
i.e., that the Hadley cell is roughly 4 times as strong as the Ferrell cell**



Note that we are able to prove the existence of Hadley and Ferrel cells and to reliably estimate their magnitudes without using observations of $[v]$. Until the 1980s, direct estimates of $[v]$ were unreliable.